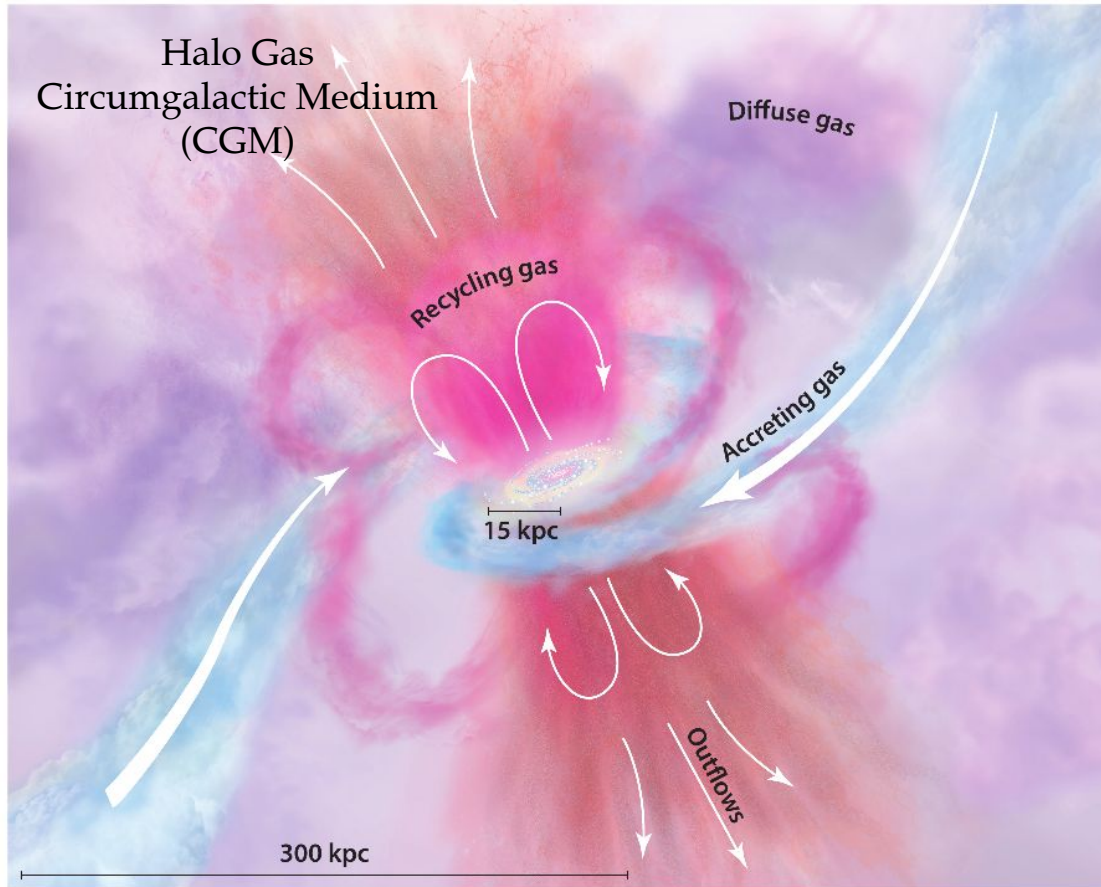


The Galaxy-Halo Gas Connection



Three Main Messages

1. (Cool) Halo gas (kinematics) traces the galaxy stellar mass.
2. Highly-ionized (warm/hot?) halo gas traces star-forming galaxies.
3. Environment suppresses halo gas in satellite galaxies.
- (4. Quasars show the extrema of Cool Halo gas)

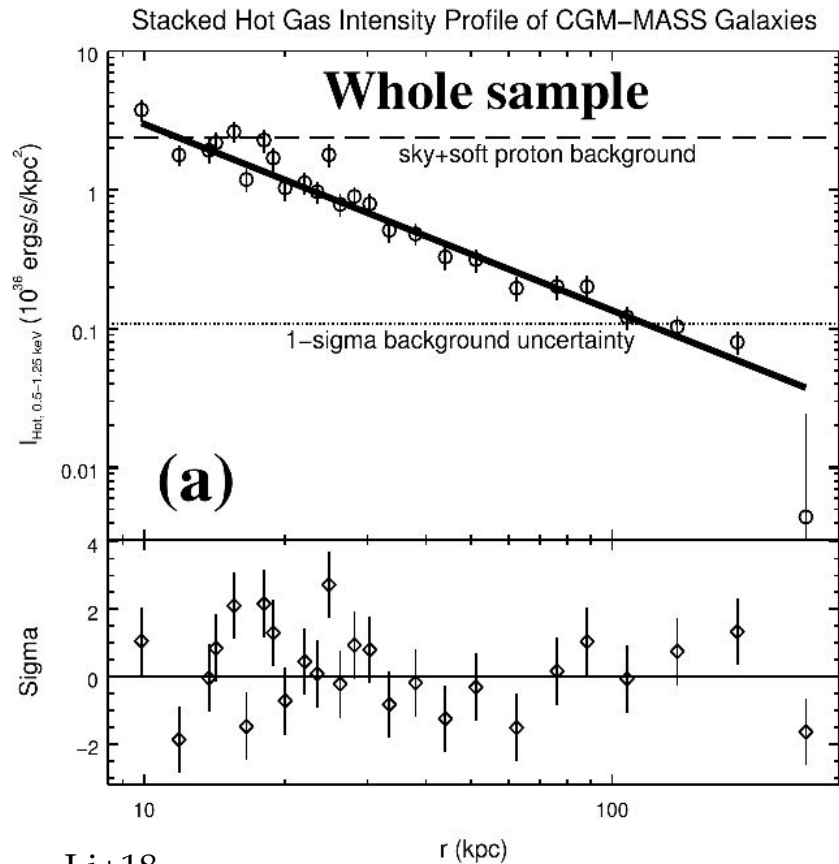
Caveats

- $z < 1$ (primarily)
- Hot gas is nearly unconstrained
- Sample sizes are modest [10's, not even 100's]

X-ray Emission: Too Diffuse/Soft for Galaxy Halos

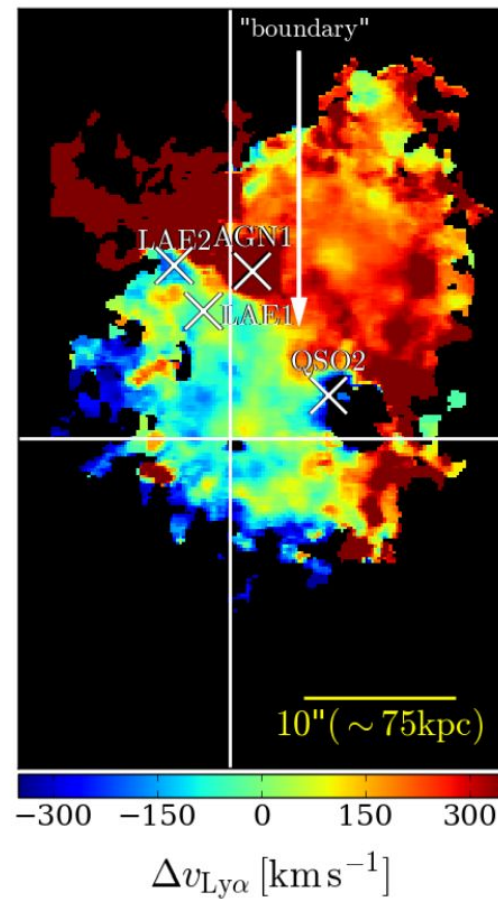
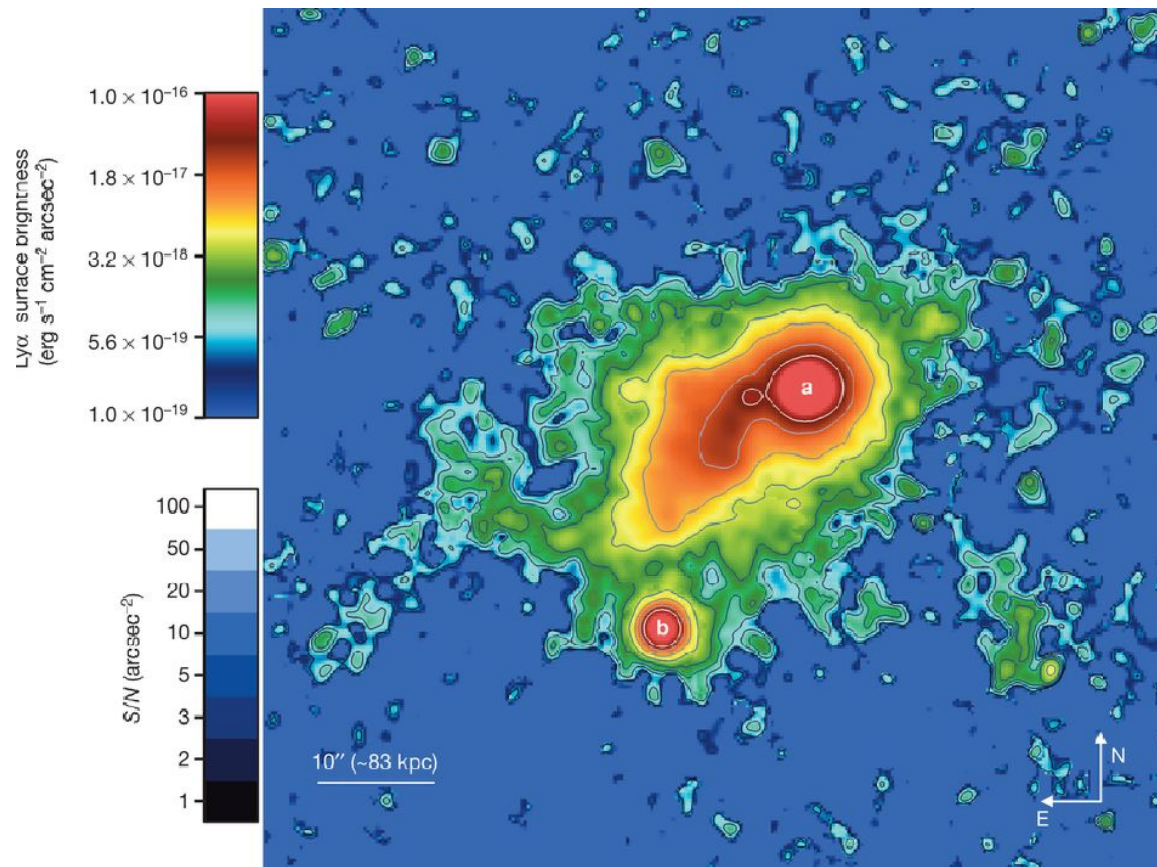


Dali 2018

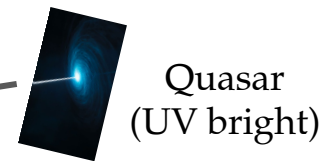
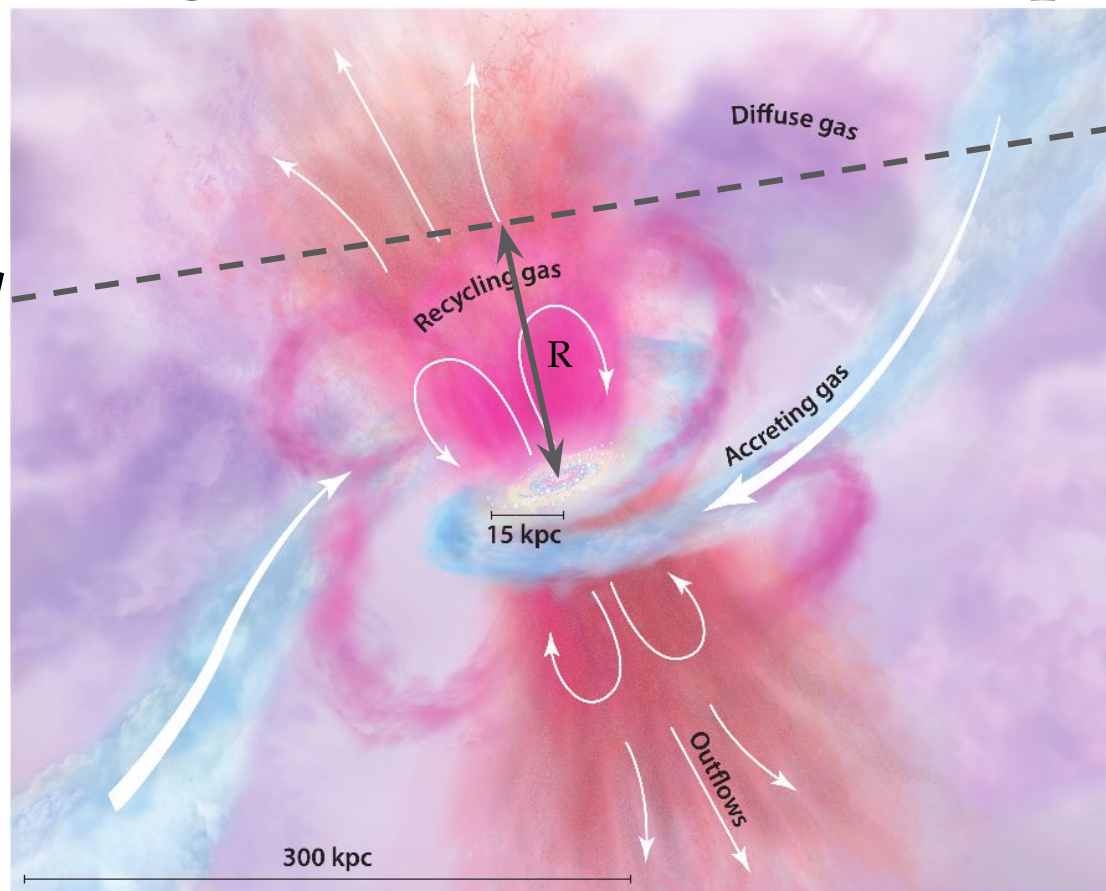
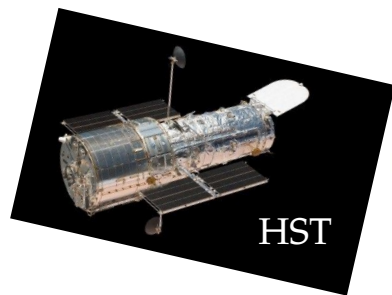


Li+18

$\text{Ly}\alpha$ Emission: Progress with Cool HI Gas

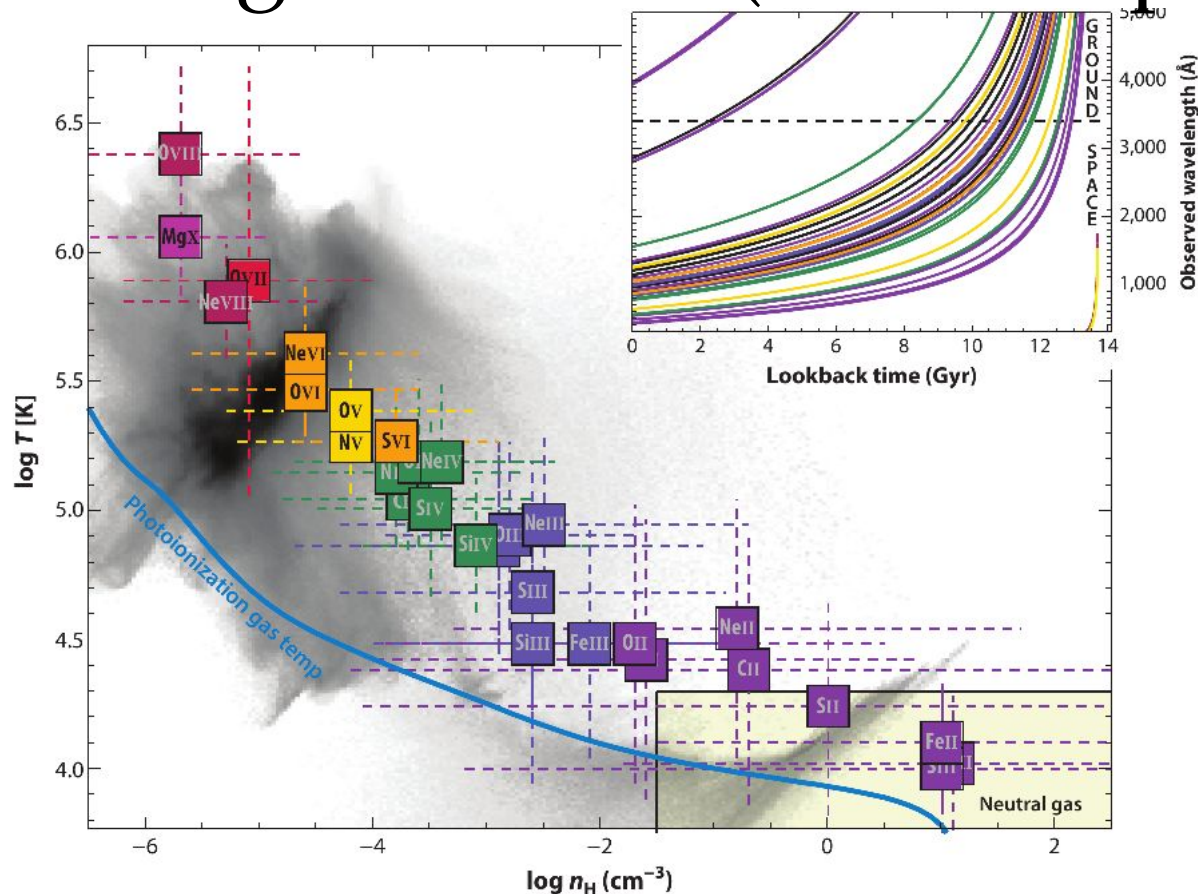


Probing Halo Gas in Absorption



Tumlinson+17

Diagnosing Halo Gas (in Absorption)

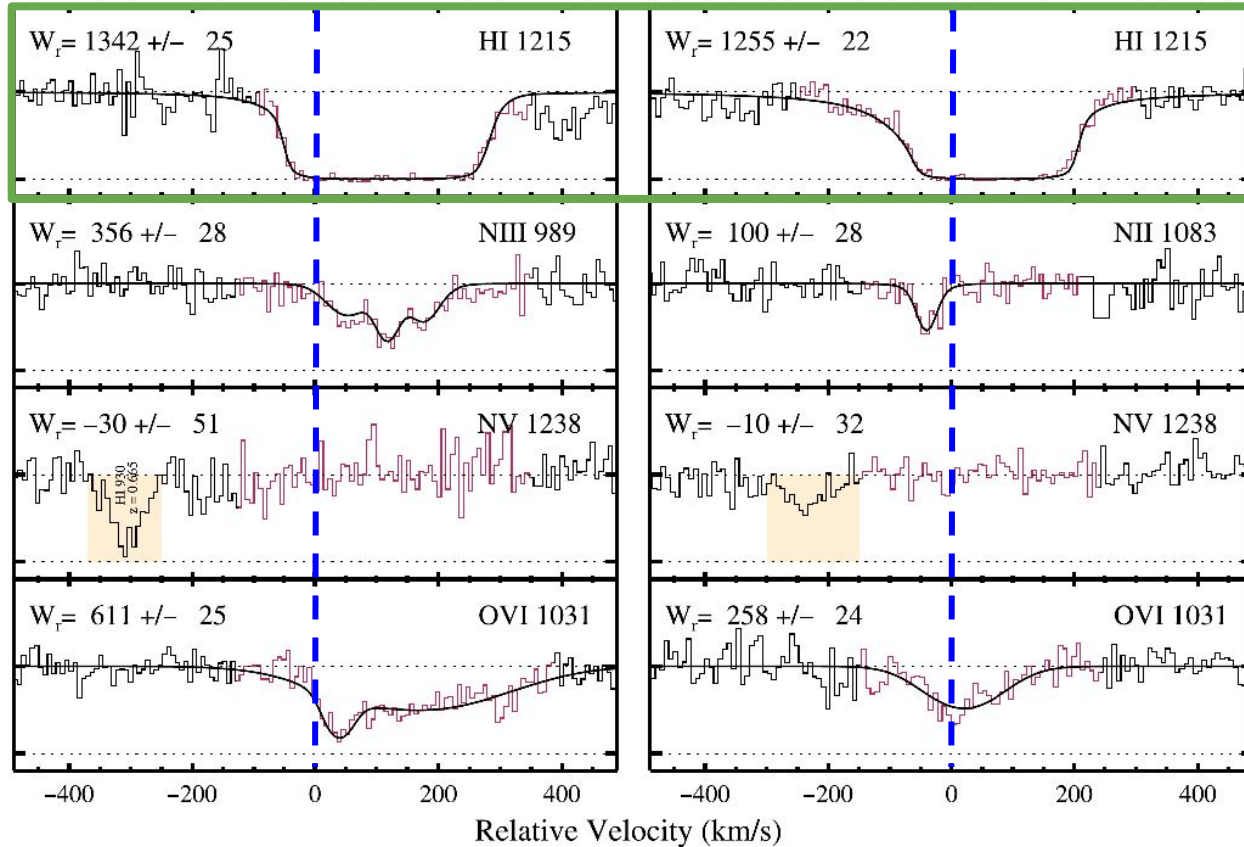


Tumlinson+17

Halo Gas Detected

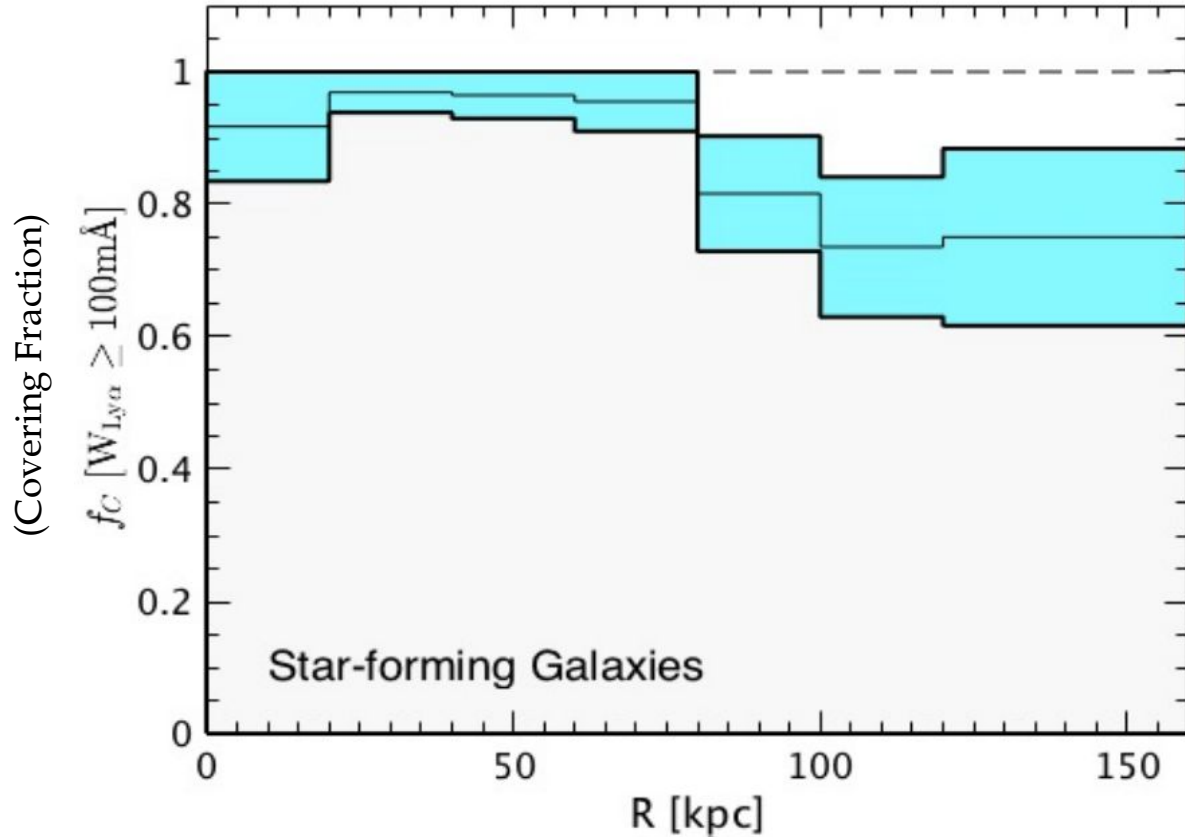
R=23 kpc $z_{\text{gal}} = 0.252$ J1016+4706: 274_6

$z_{\text{gal}} = 0.192$ J1330+2813: 289_28 R=87 kpc



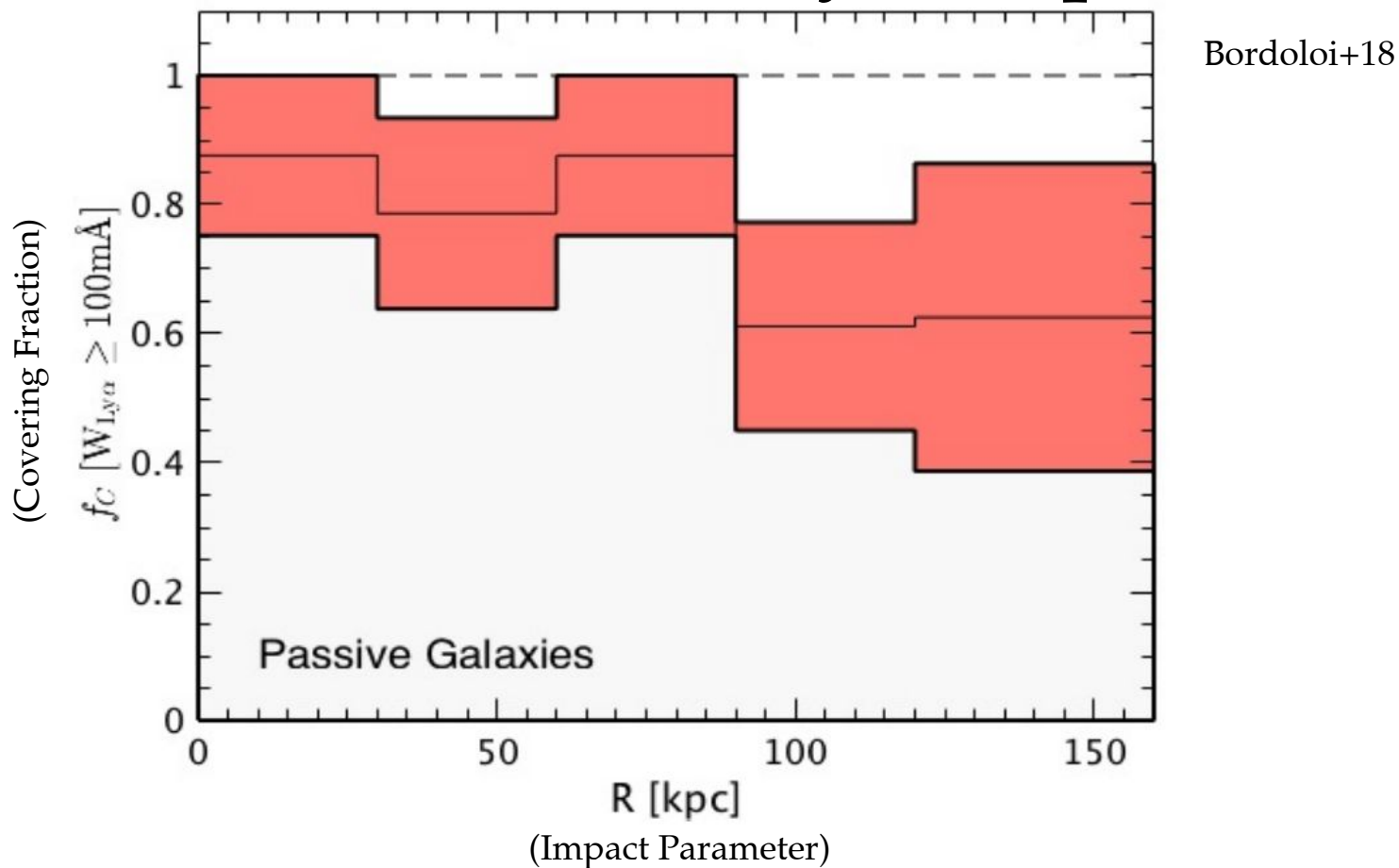
Werk+16

Cool Halo Gas is nearly Ubiquitous

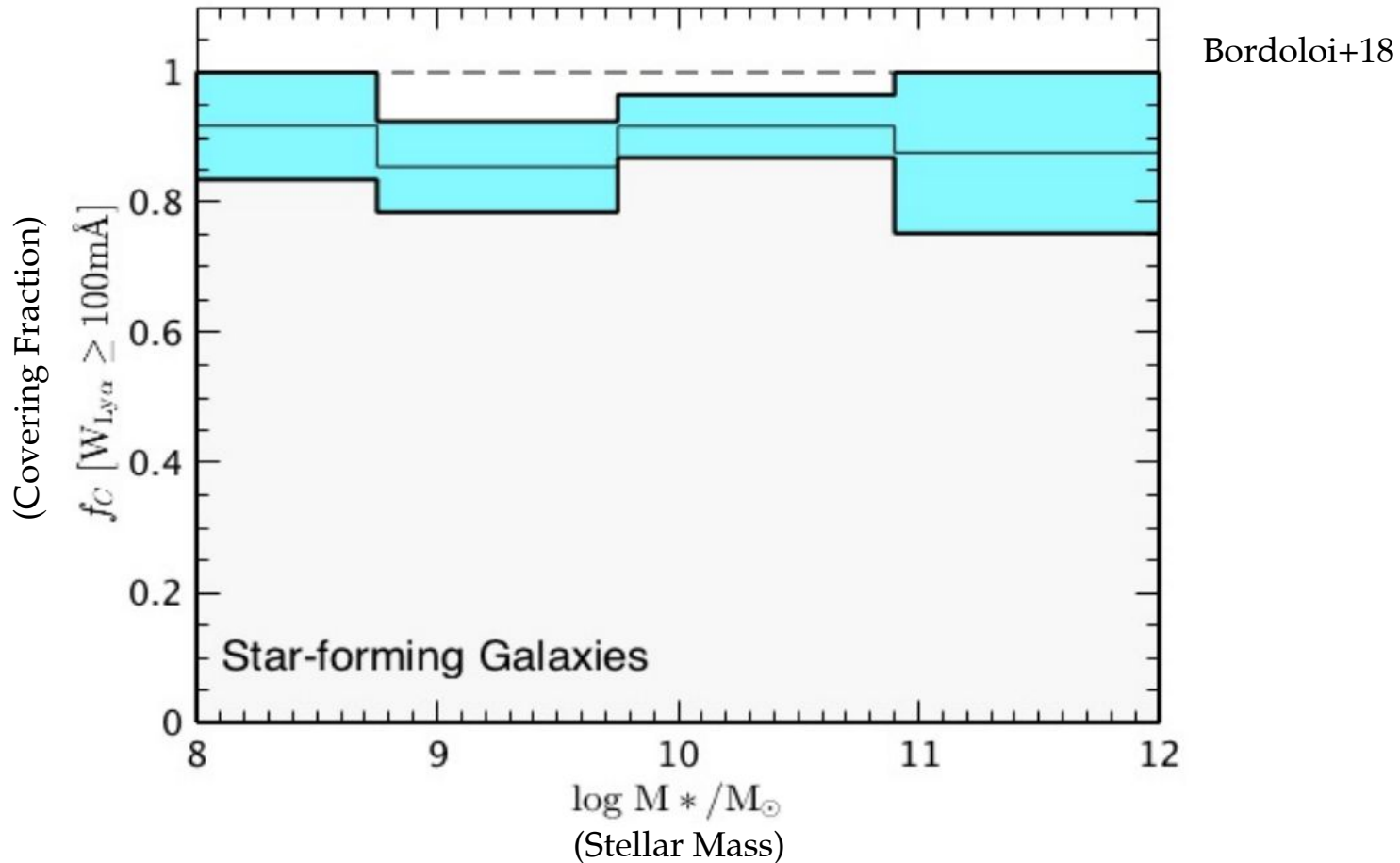


X+11
Tumlinson+13
Bordoloi+16,18

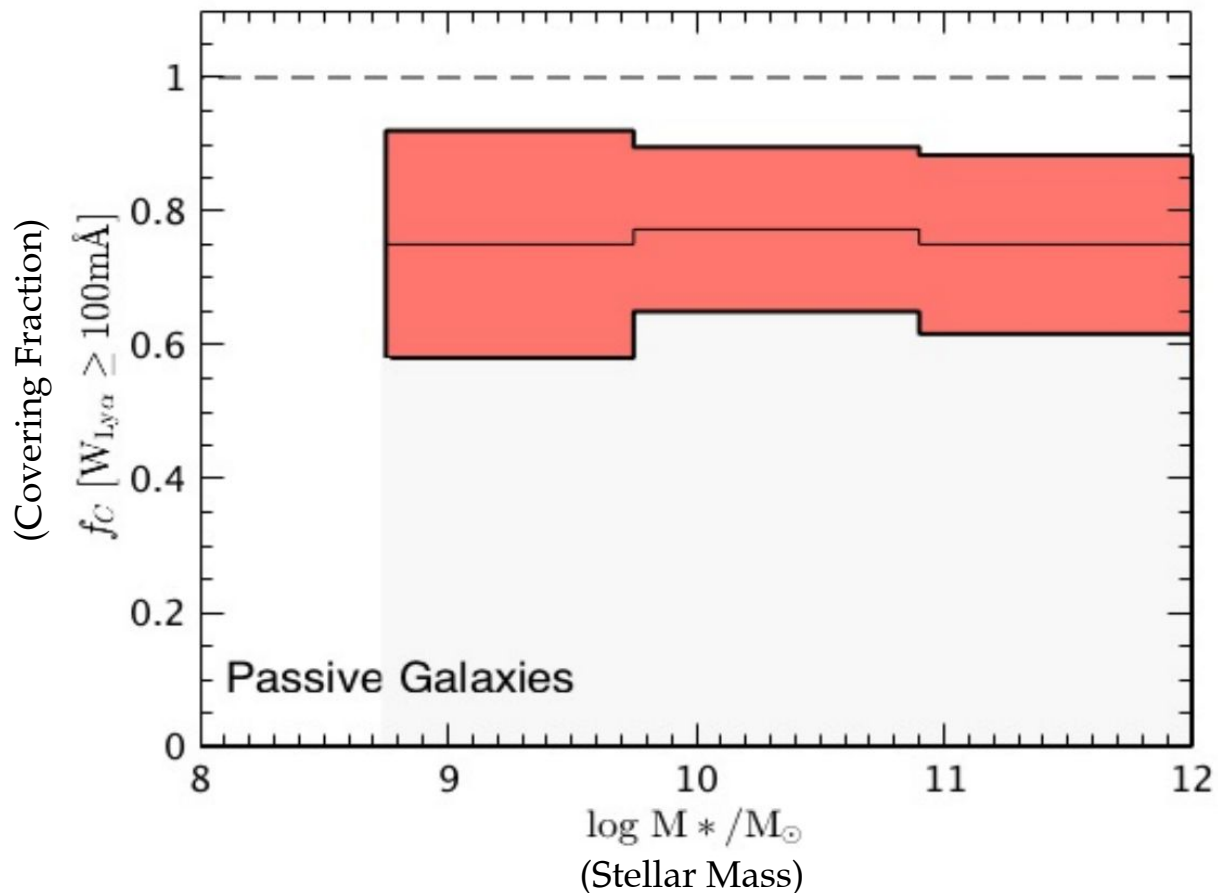
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Cool Halo Gas is nearly Ubiquitous

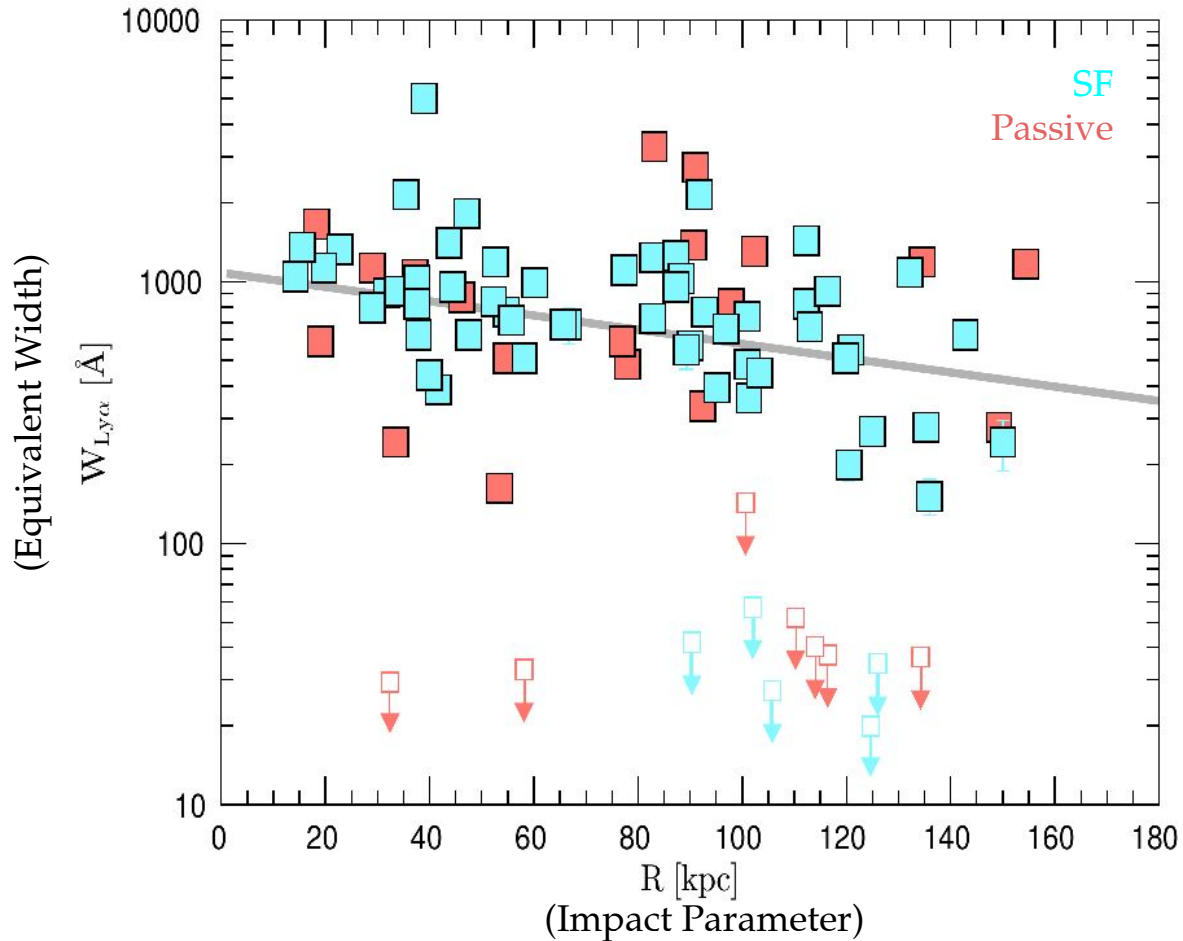


Cool Halo Gas is nearly Ubiquitous

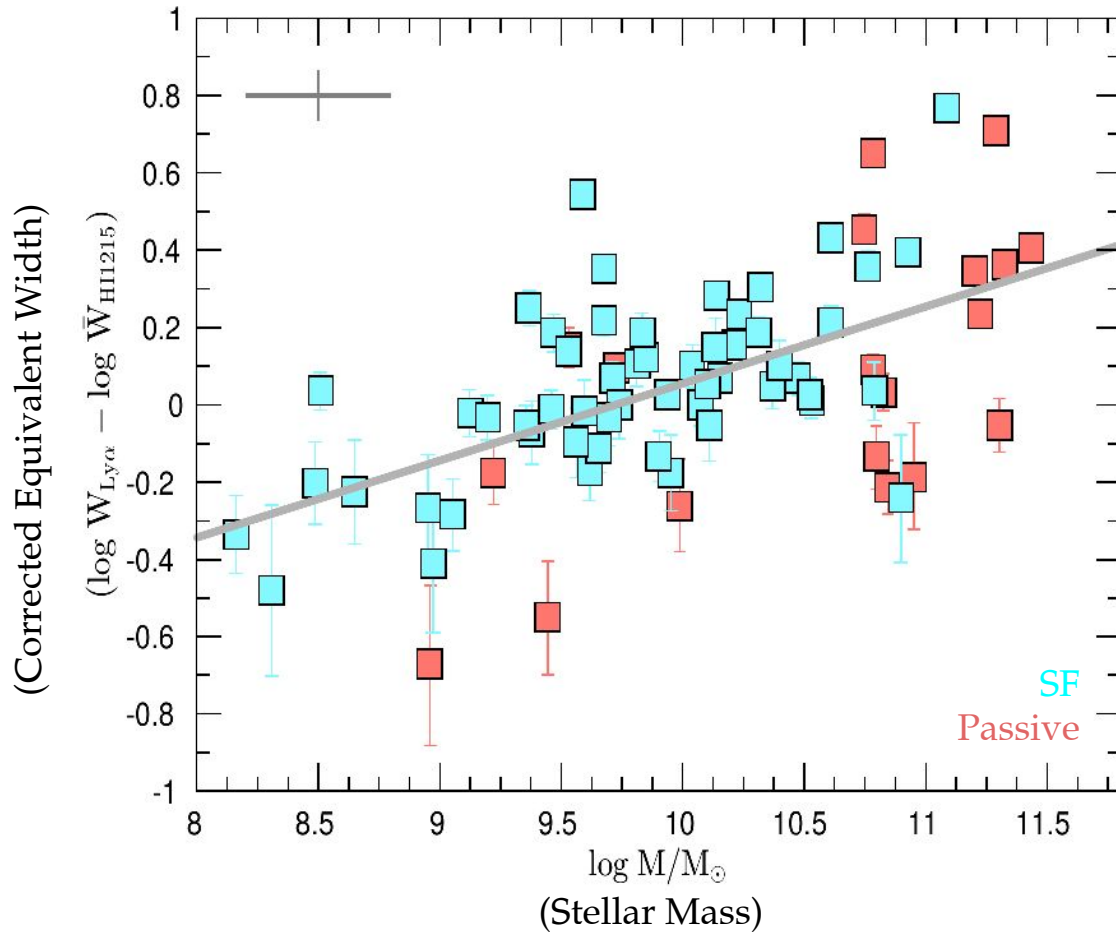


Bordoloi+18

Cool Halo Gas scales with R



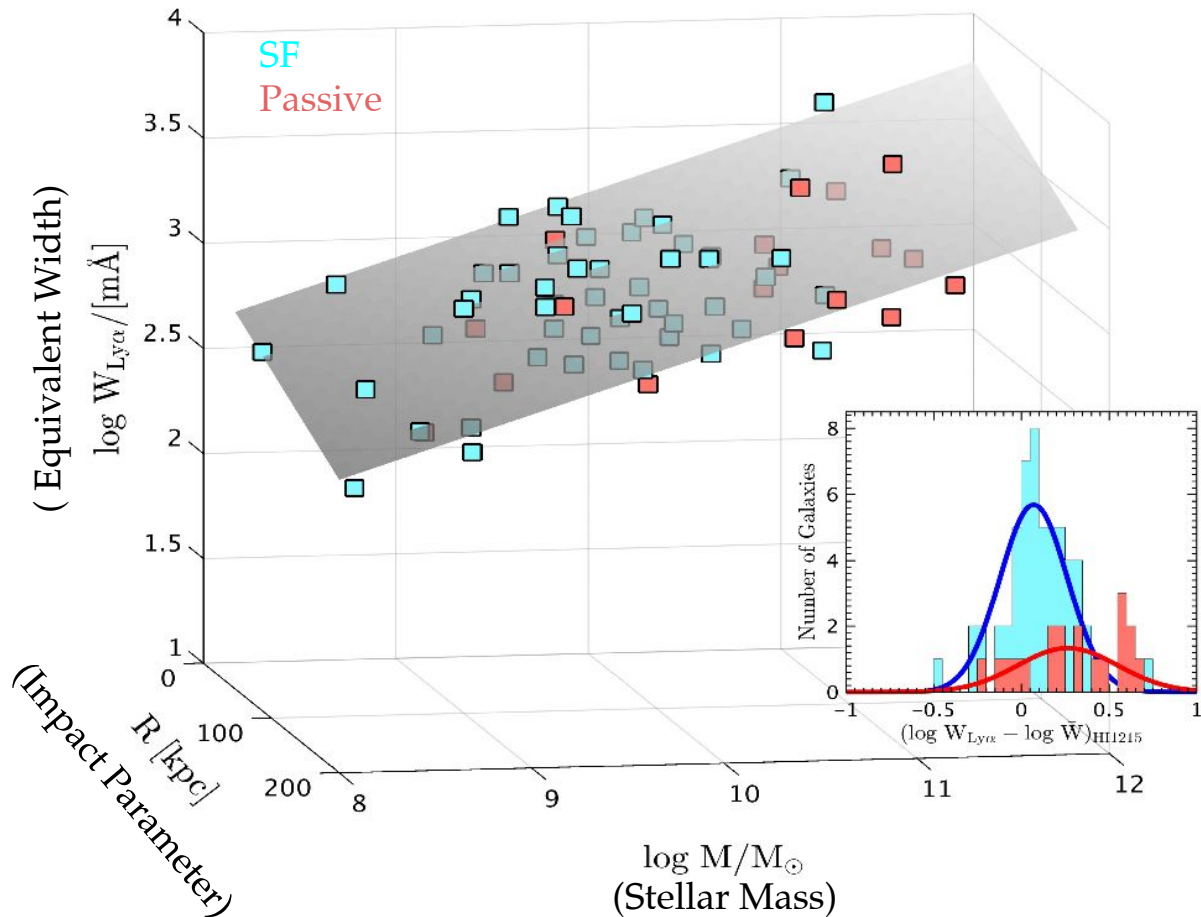
Cool Halo Gas scales with M^*



Bordoloi+18

Cool Halo Gas: Fundamental Plane

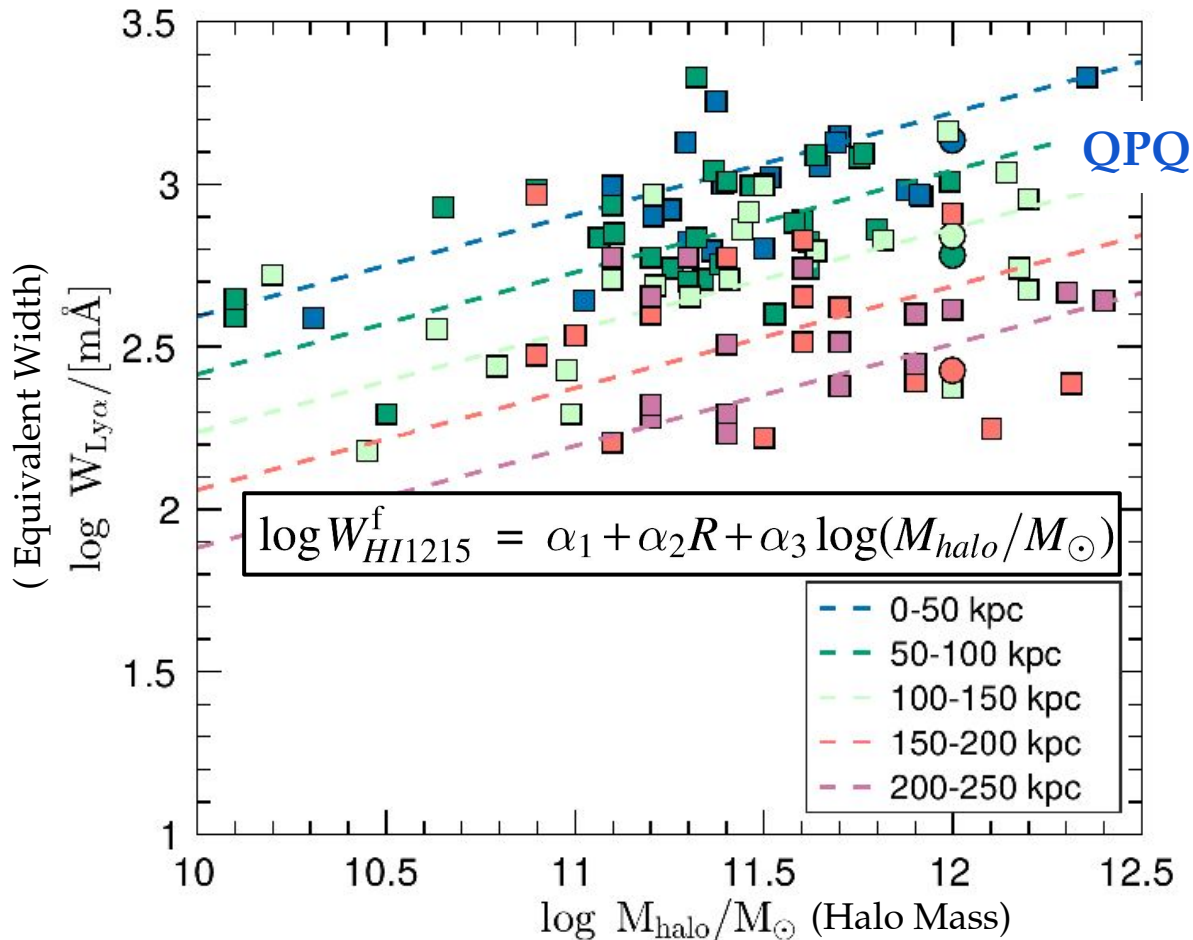
Bordoloi+18



Cool Halo Gas Traces the Halo Mass

LRGs*

Bordoloi+18



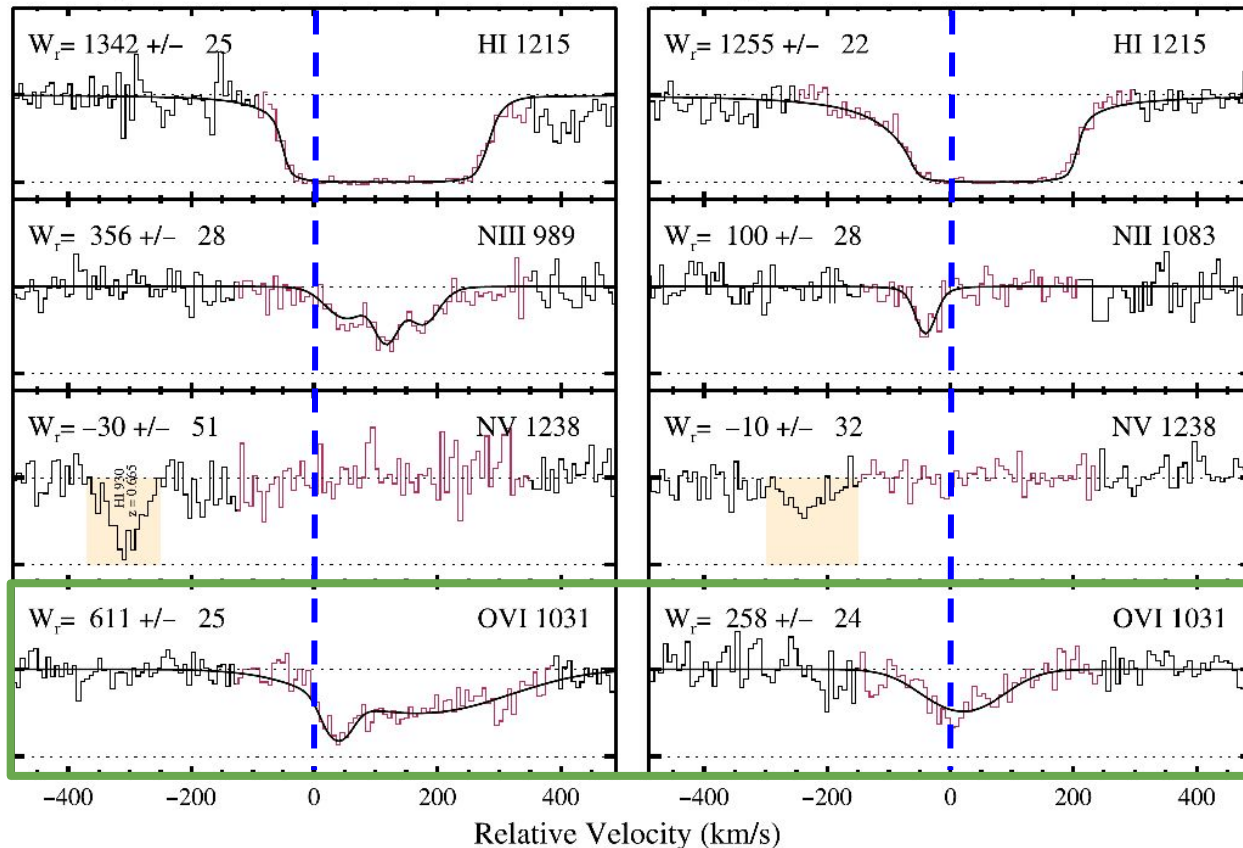
Dali 2018

Halo Gas Detected

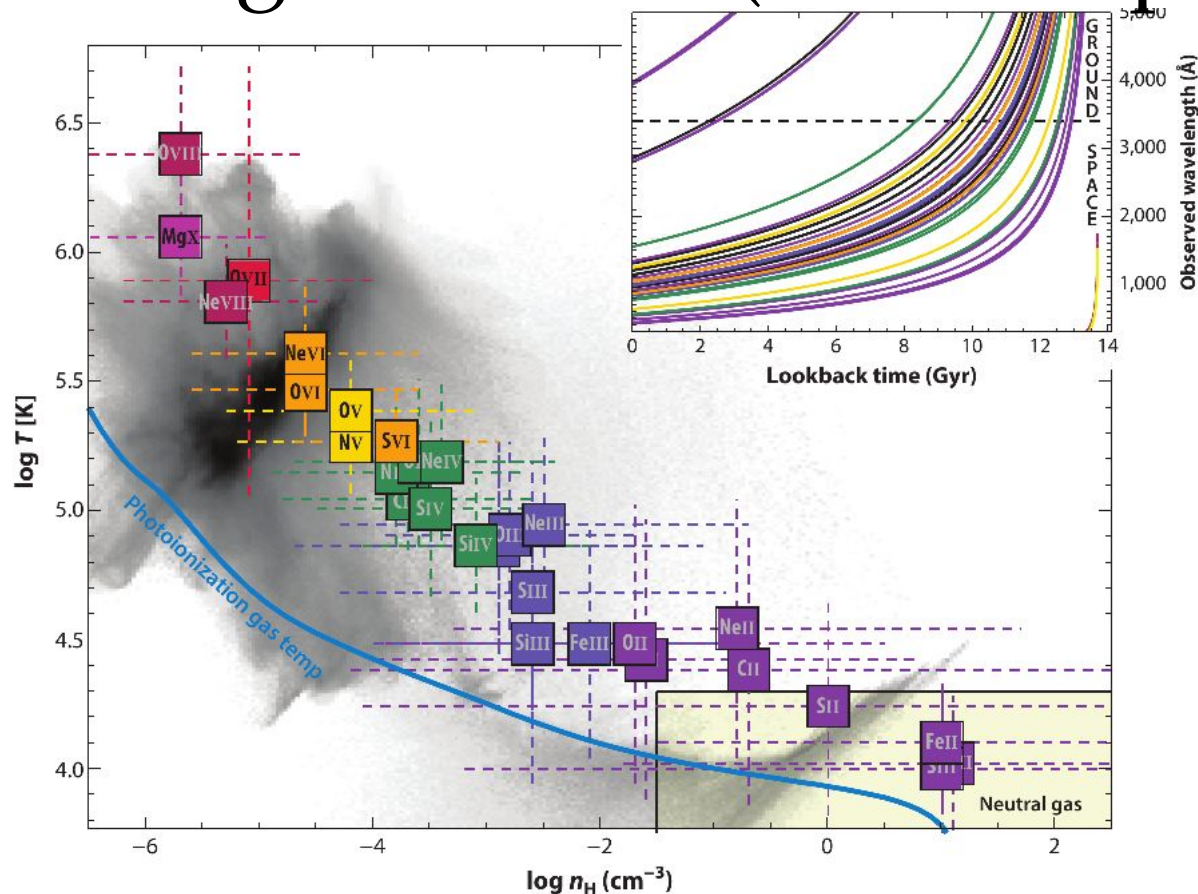
R=23 kpc $z_{\text{gal}} = 0.252$ J1016+4706: 274_6

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Werk+16

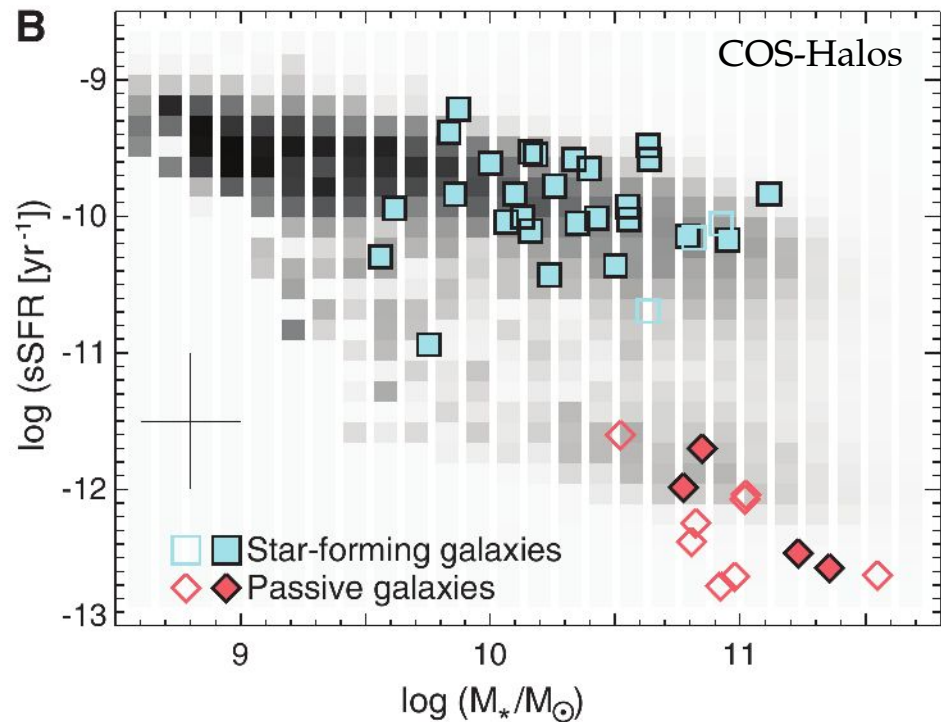


Diagnosing Halo Gas (in Absorption)



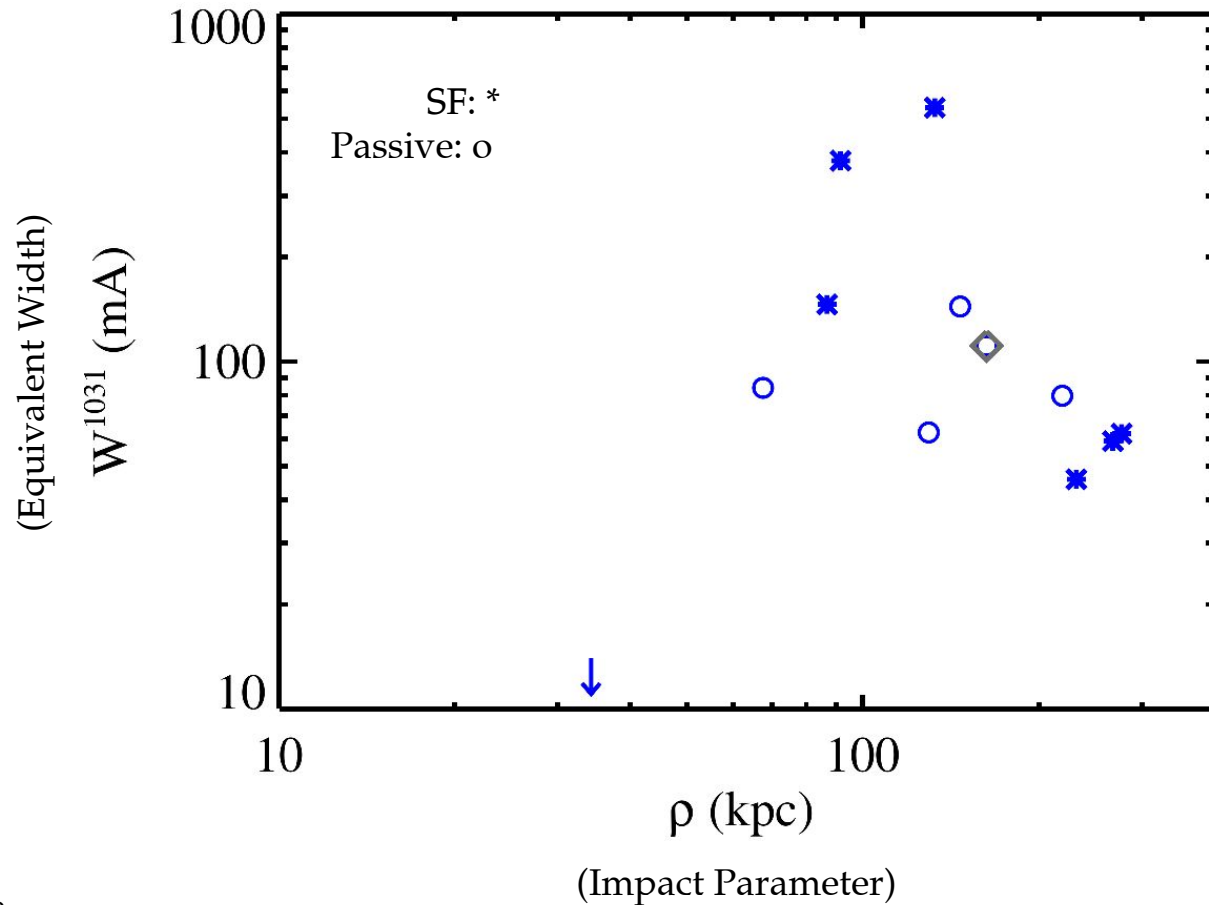
Tumlinson+17

OVI Gas traces Star-Formation in L^* Galaxies



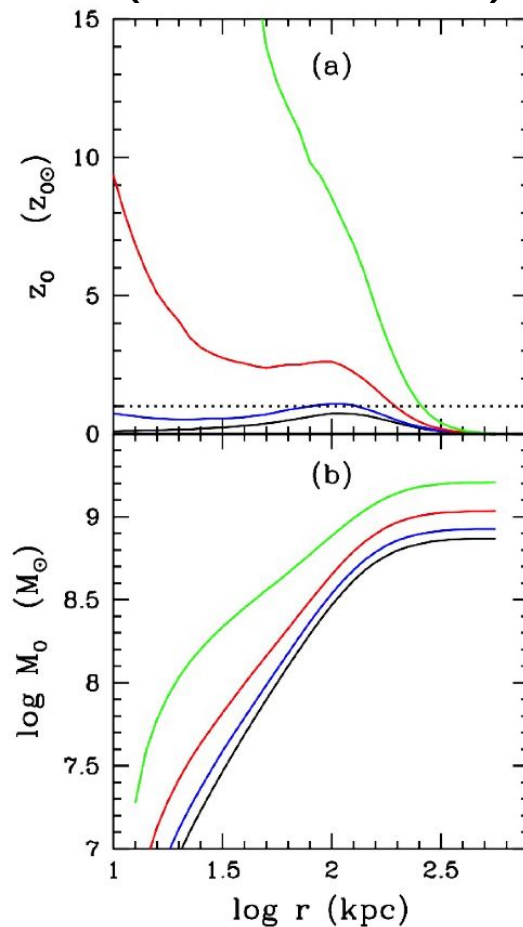
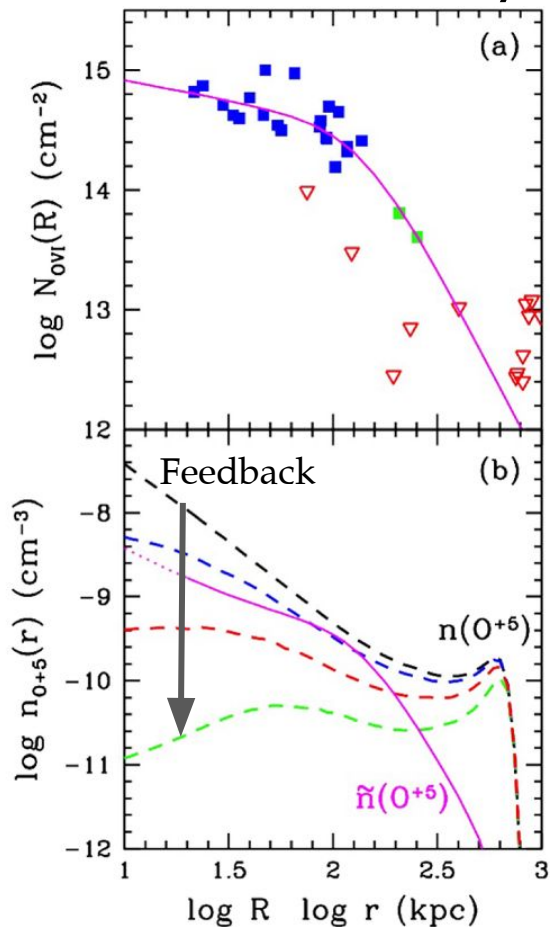
Tumlinson+11

OVI Gas traces all sub-L* Galaxies



(Note: OVI avoids Dwarf Galaxies)

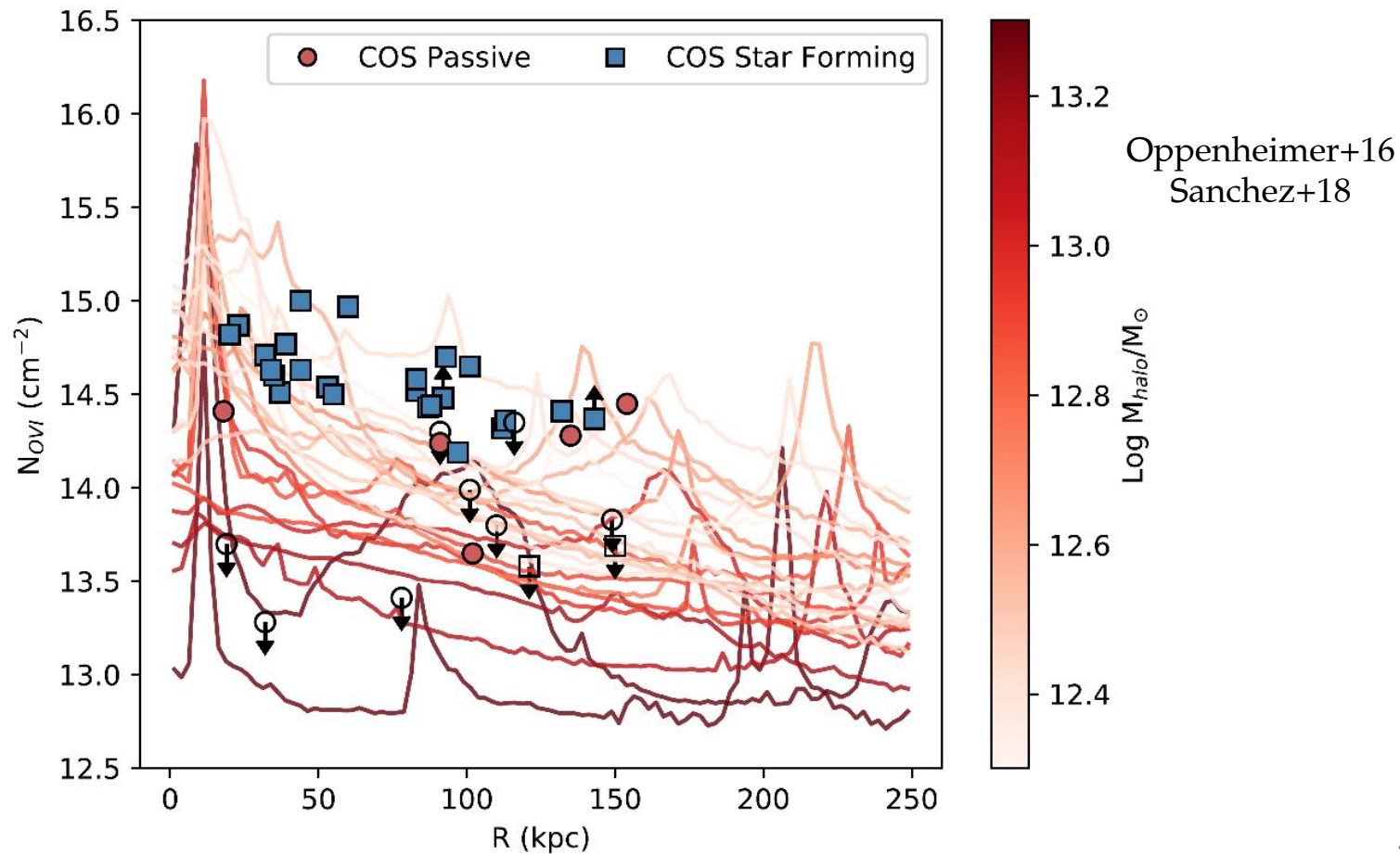
OVI traces Warm/Hot ($T \sim 10^6$ K) Halo gas?



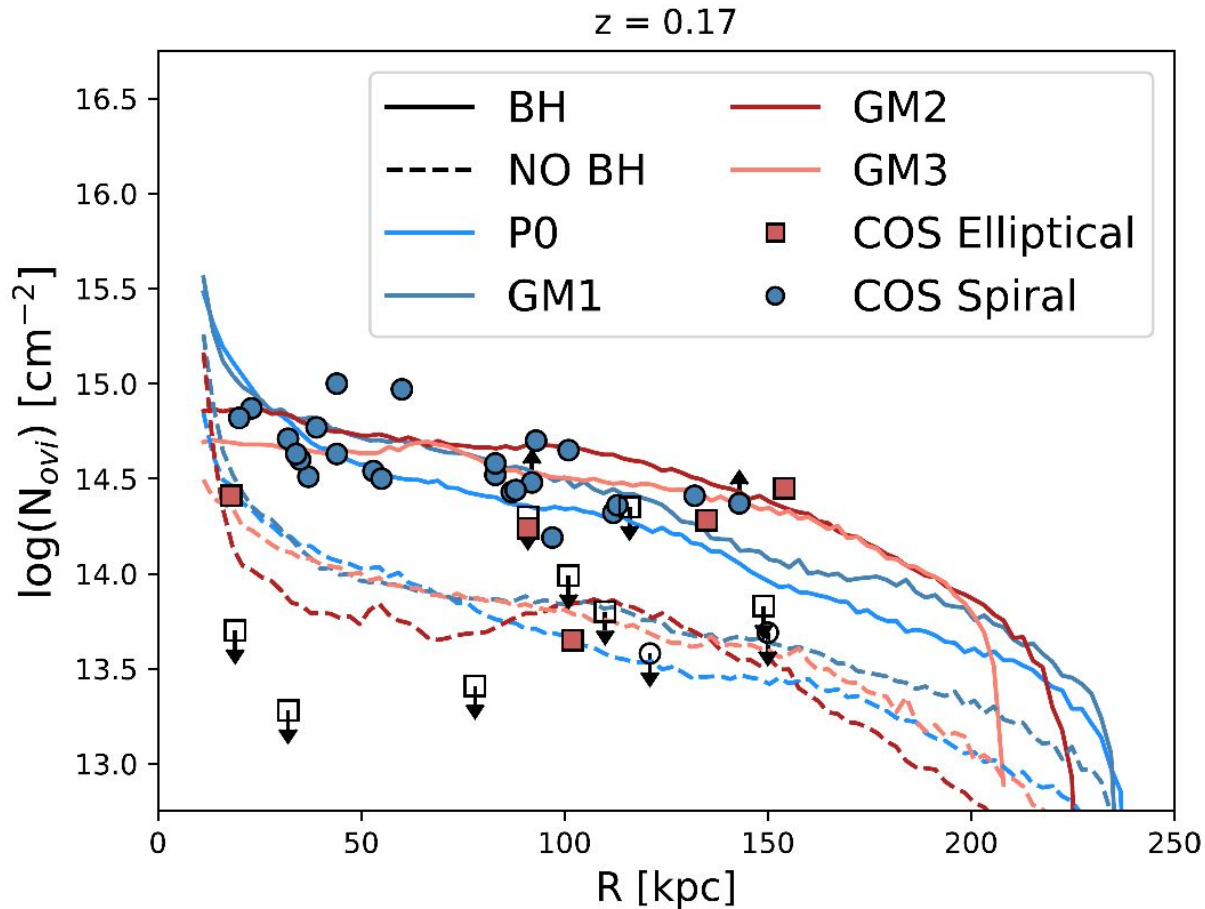
Mathews & X
(2017)

(see also
Faerman+17)

OVI Gas traces Halo Mass?



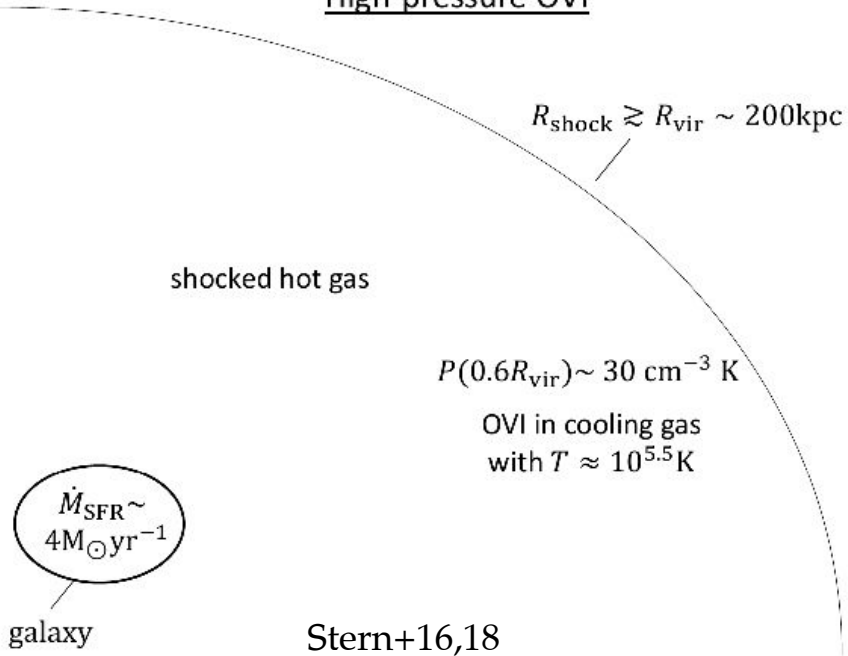
OVI Gas Needs SMBH Enrichment?



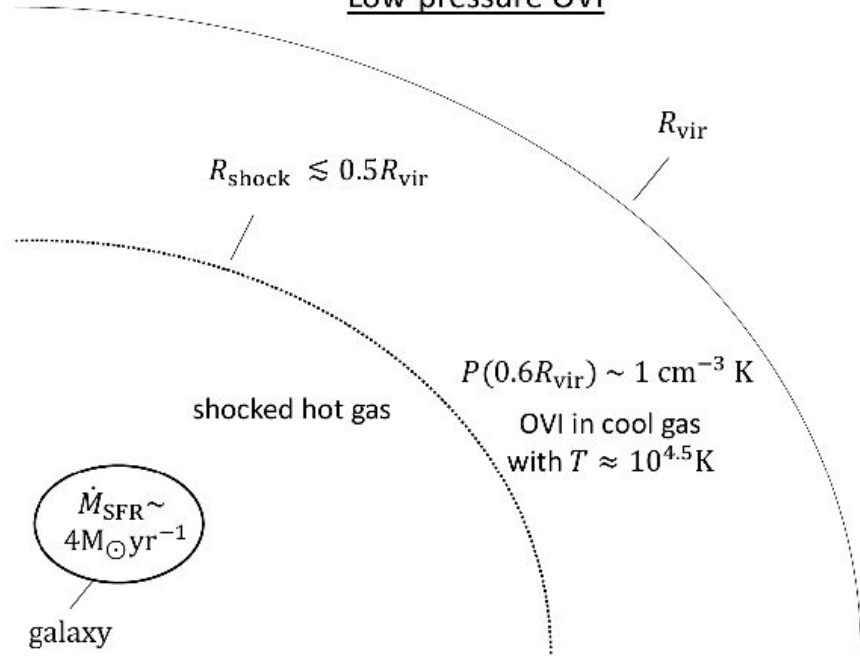
Sanchez+18

OVI Gas Traces Non-Virialized Gas?

High-pressure OVI

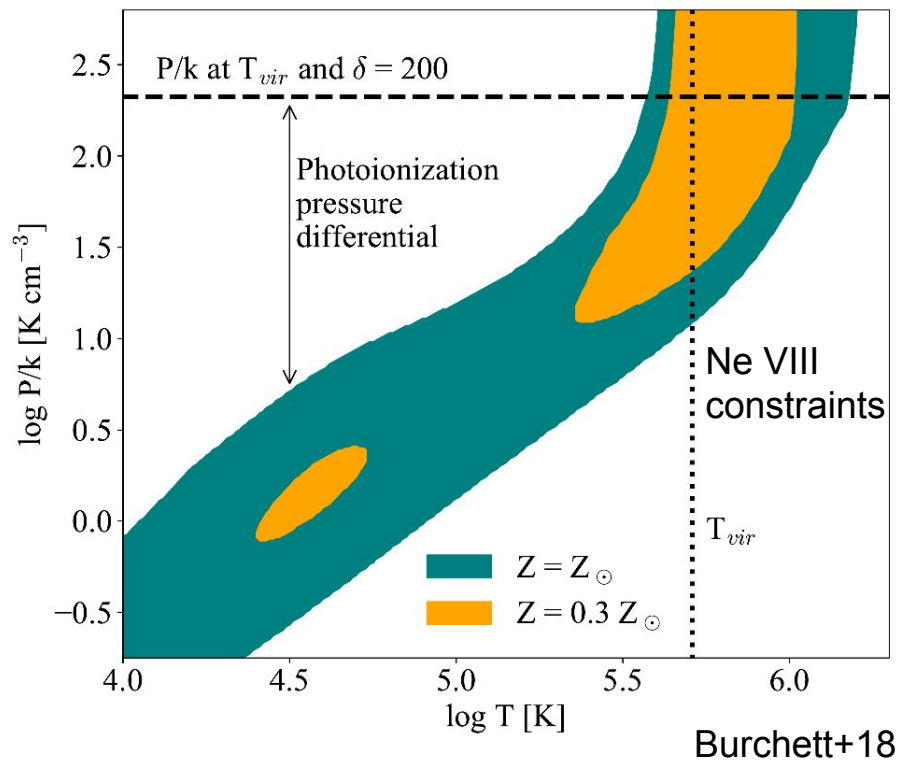
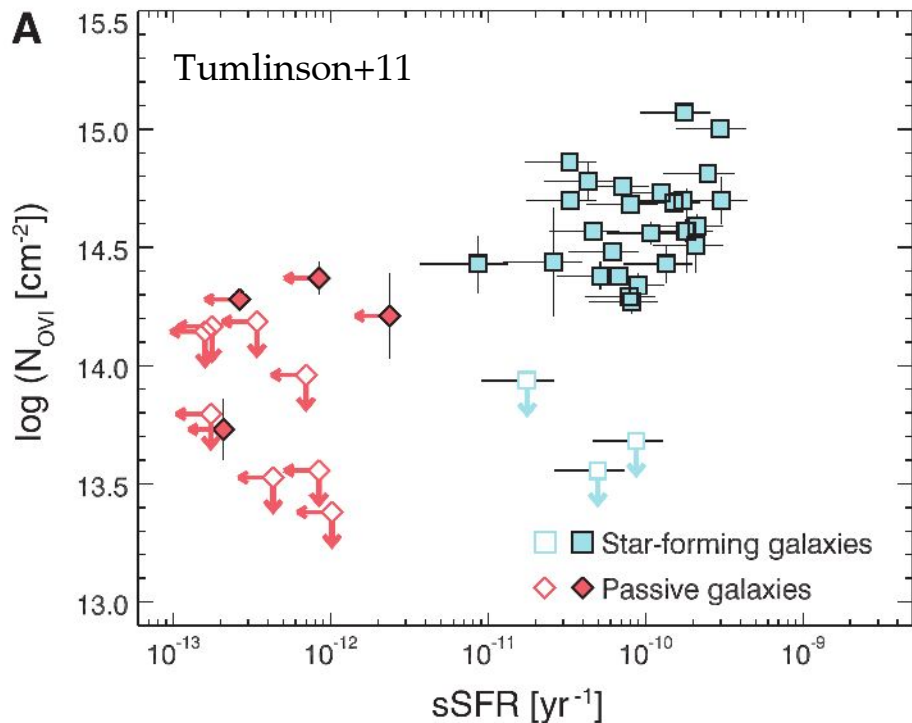


Low-pressure OVI

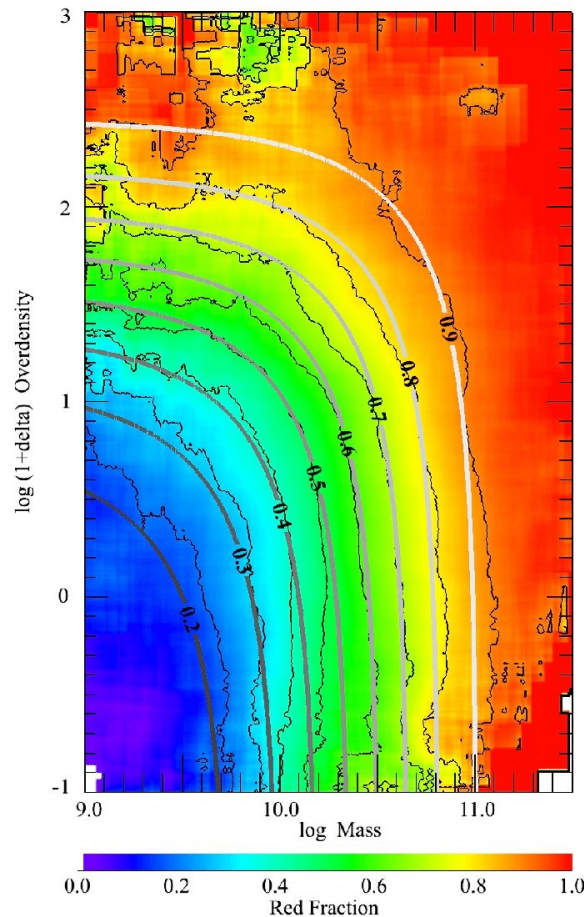
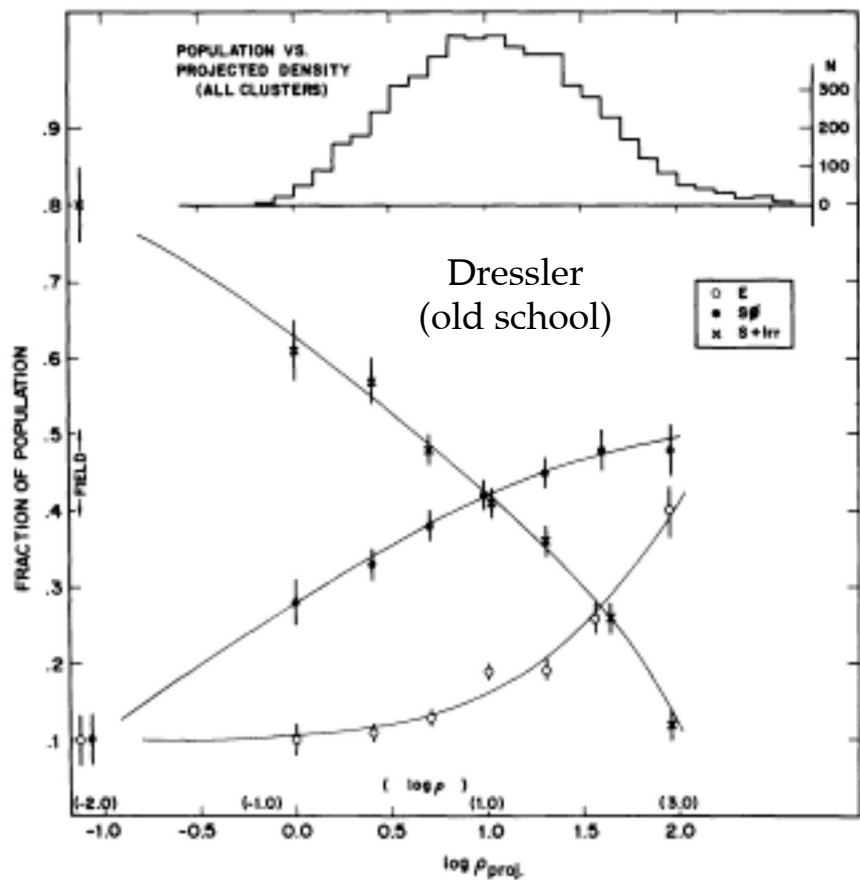


OVI Halo Knows about its Central Galaxy

But we aren't sure why...



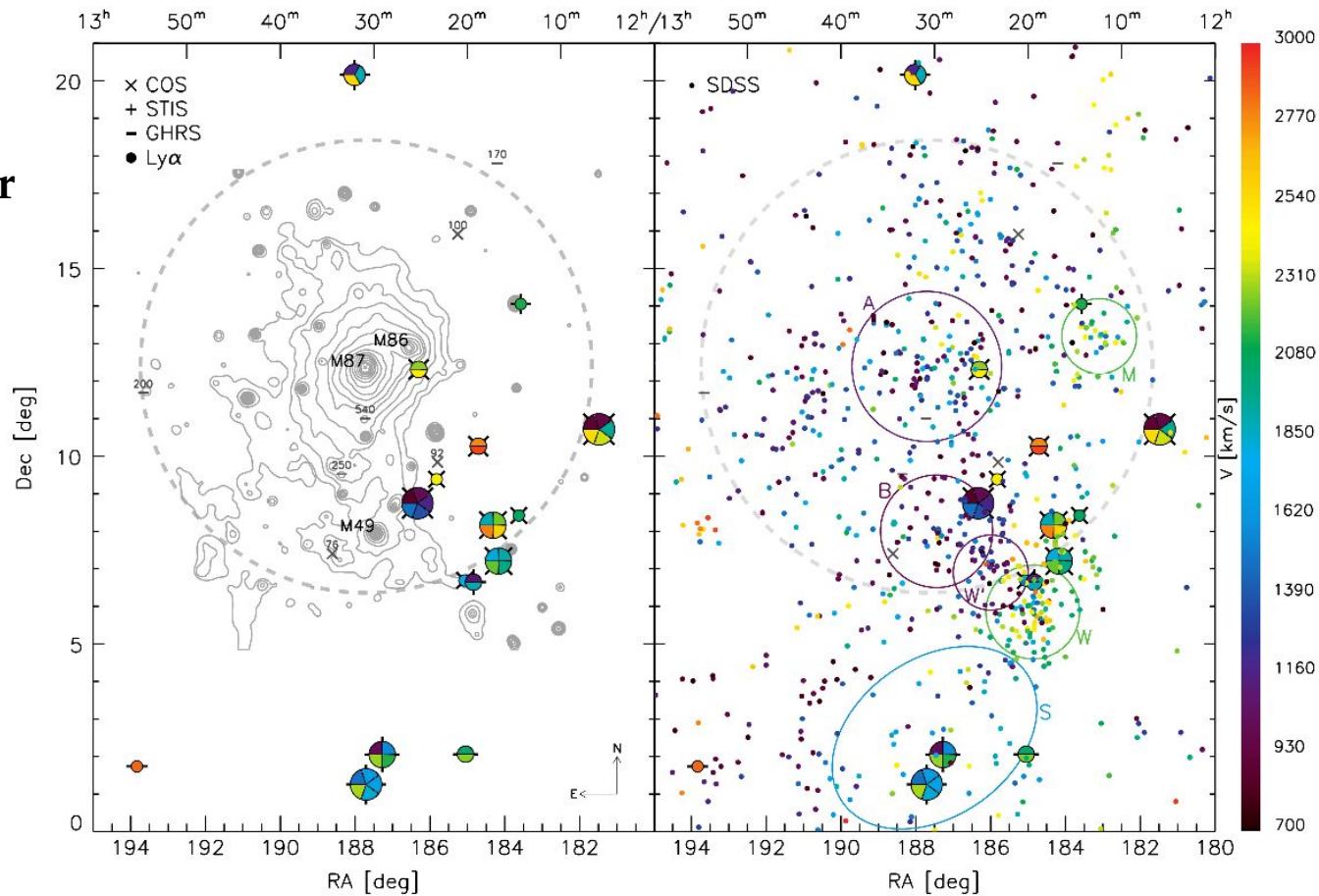
Environment Matters (for Galaxy Halos too)



Peng
(SDSS)

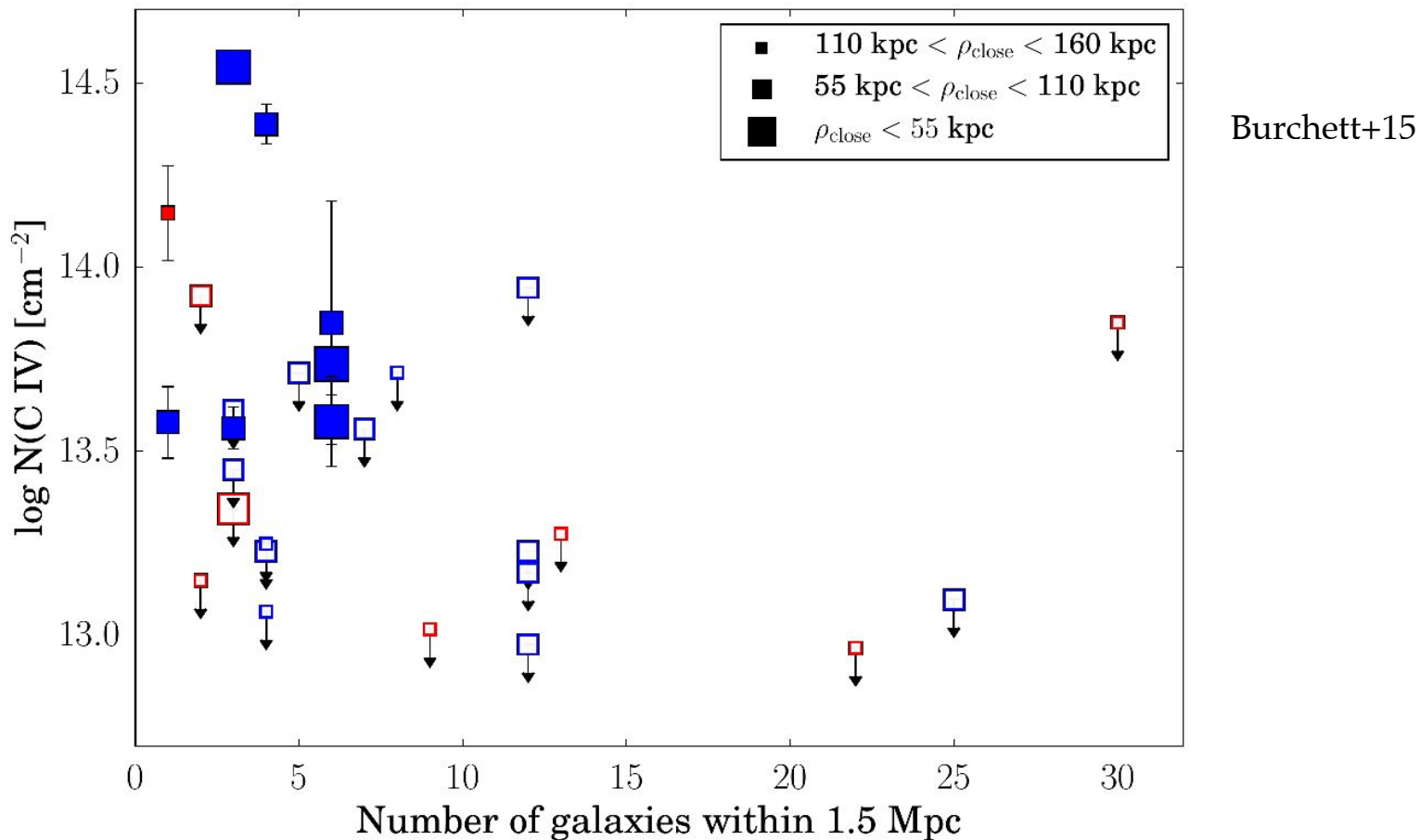
Environment Matters (Suppressed HI)

Coma Cluster

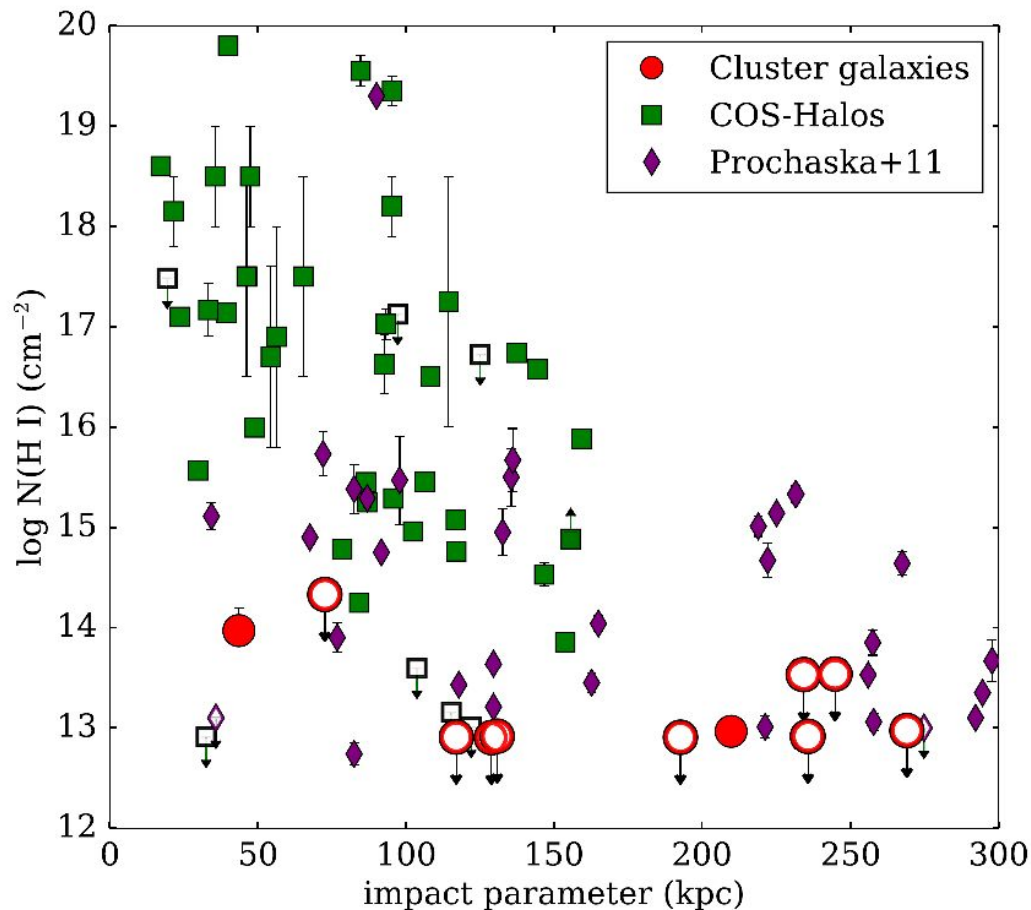


Yoon+12

Environment Matters (Suppressed CIV)



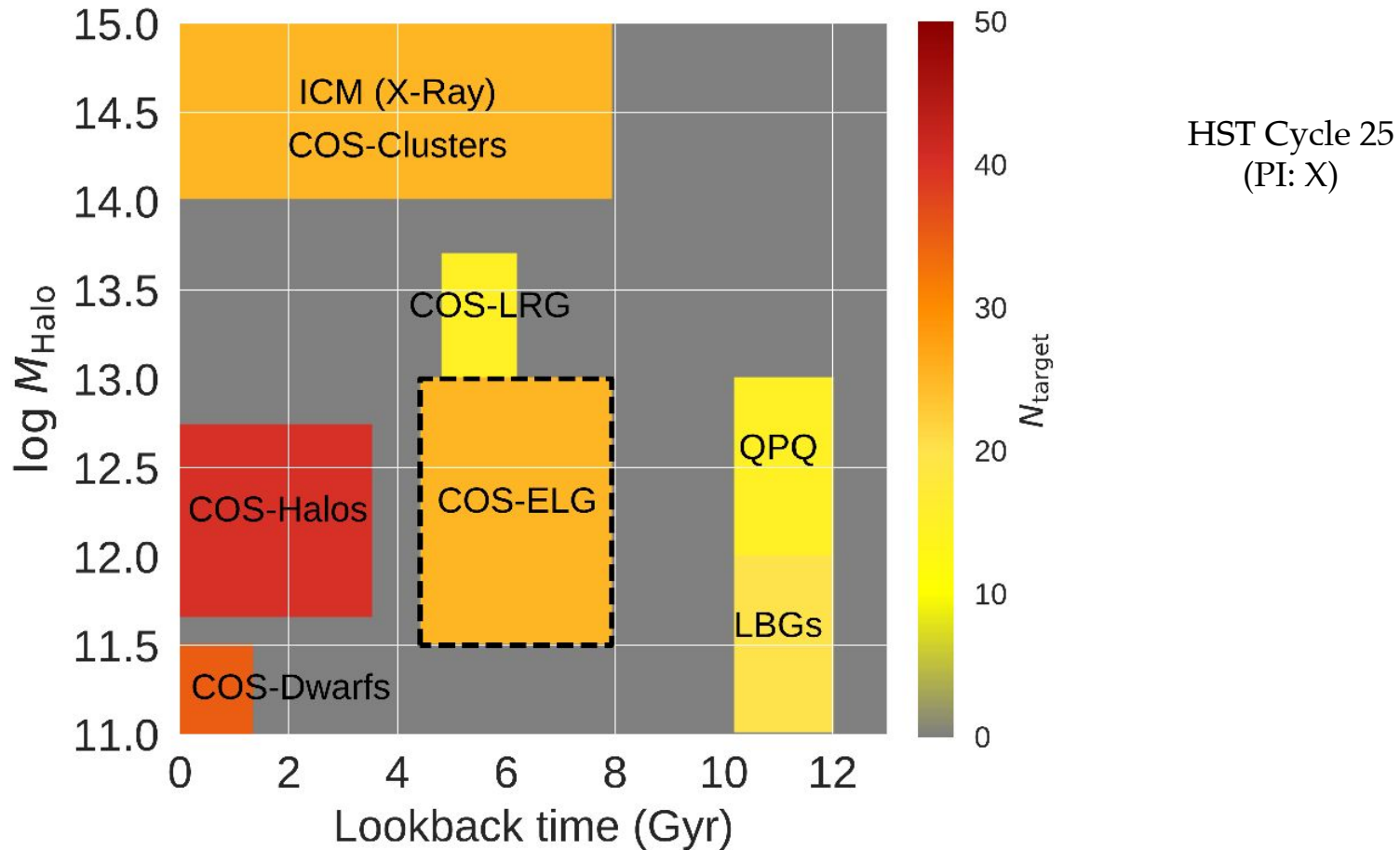
Environment Matters (Suppressed HI)



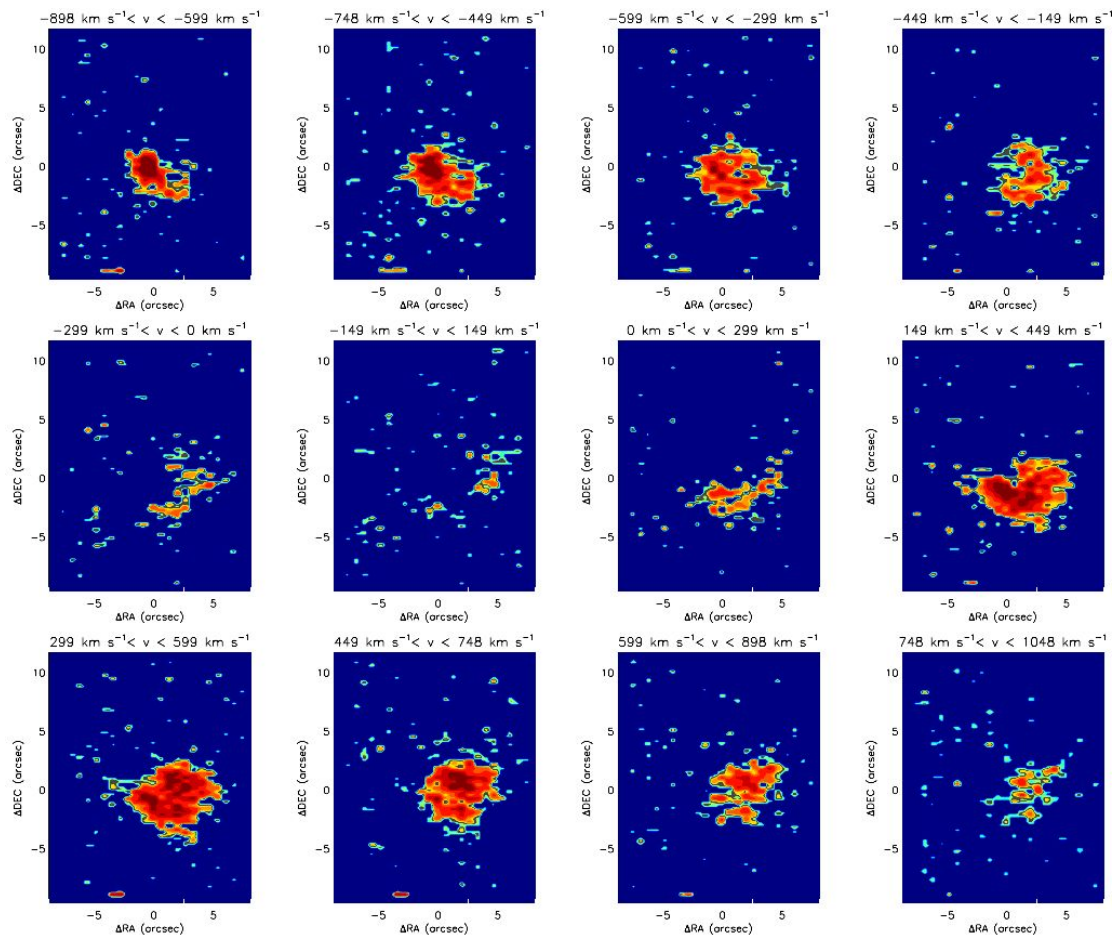
Burchett+18
Tejos+19

Is the key to
galaxy quenching
the complete
removal of a
galaxy's halo gas?

Future: CGM of ELGs

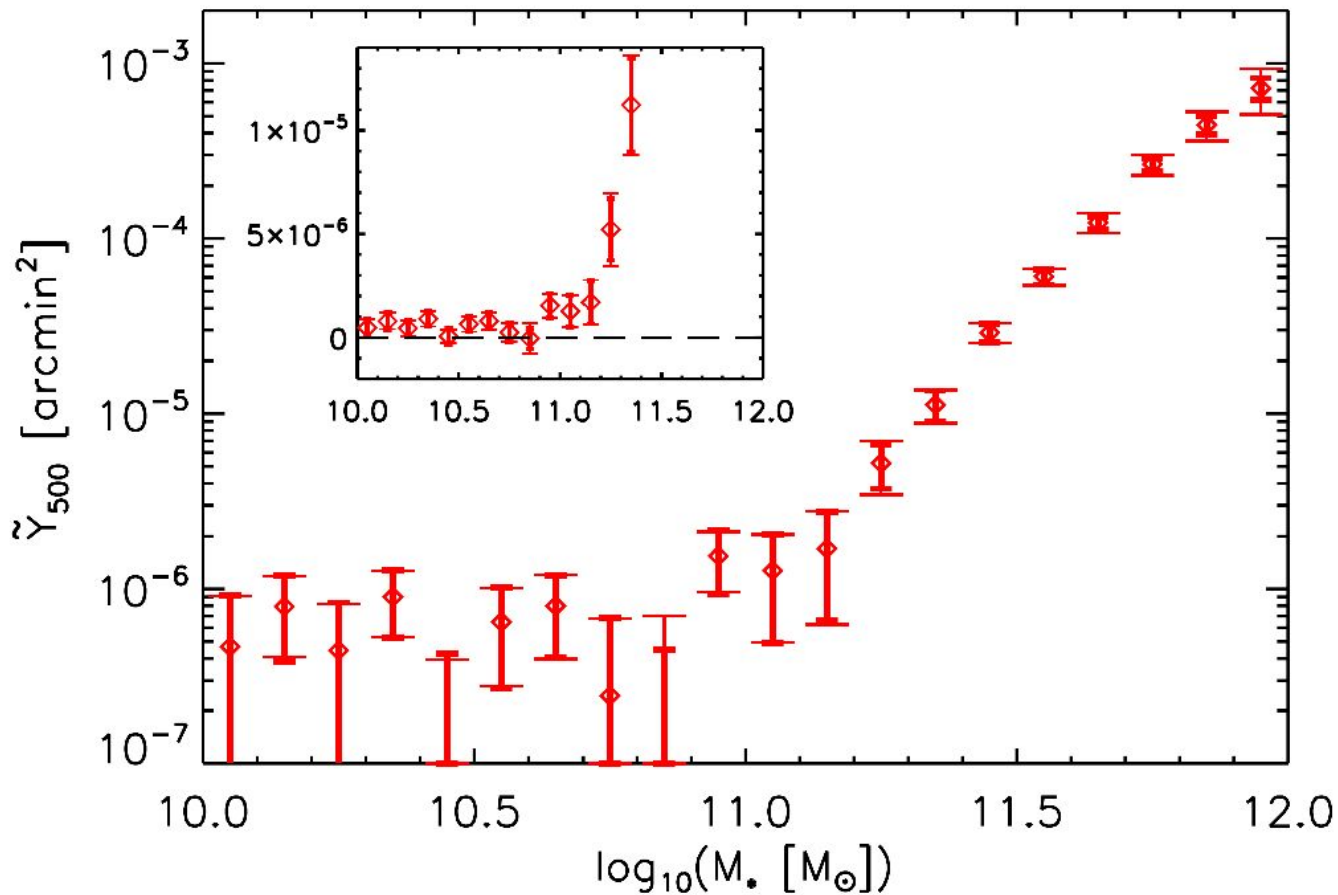


Future: CGM in Emission



Borisova+17
Cai+18

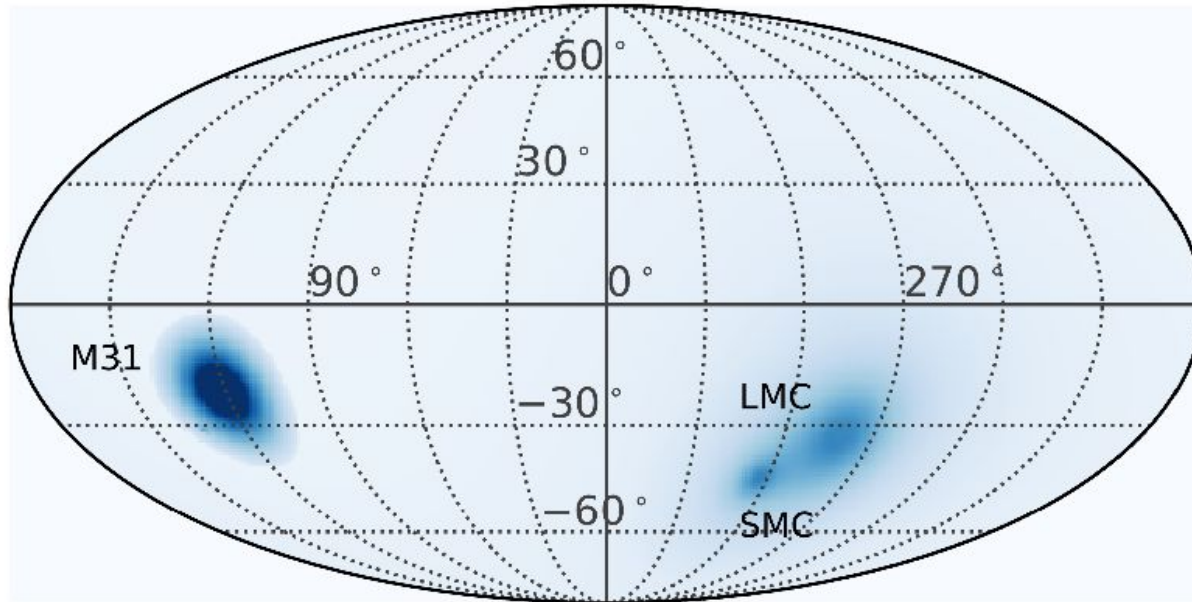
Future: SZ



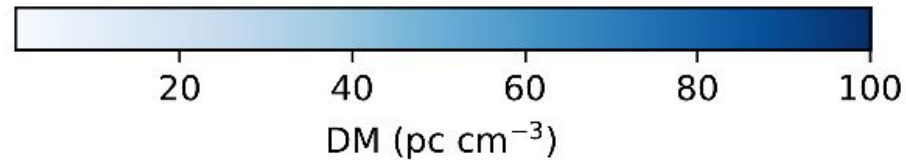
Planck

Future: FRBs

DM (M31/M33/Magellanic Clouds)

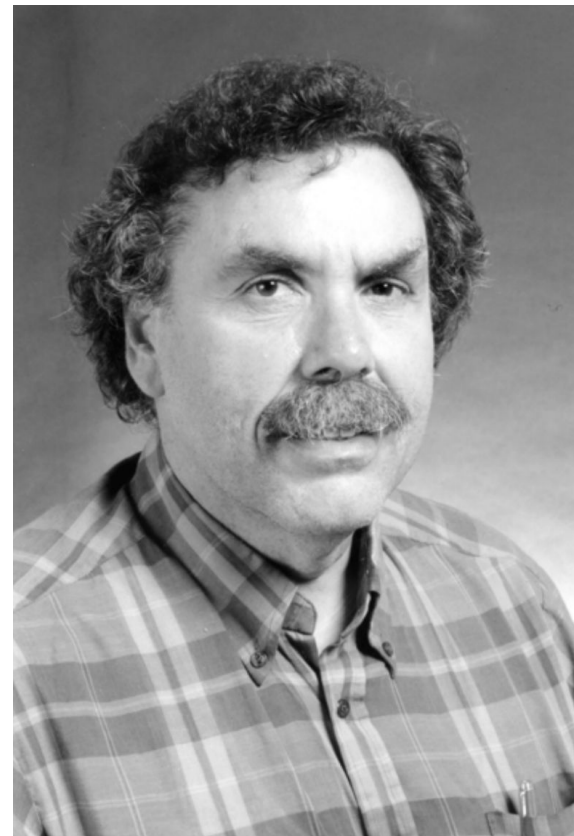
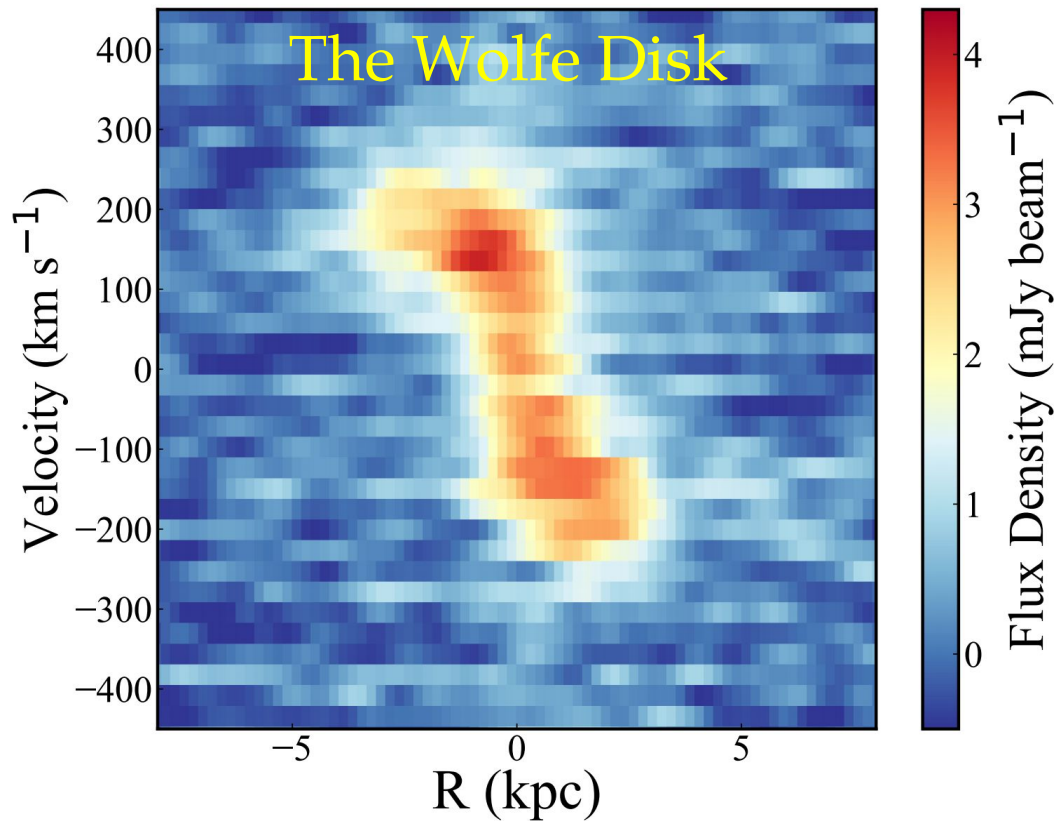


XYZ



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Cool Halo Gas Surrounding Quasars

